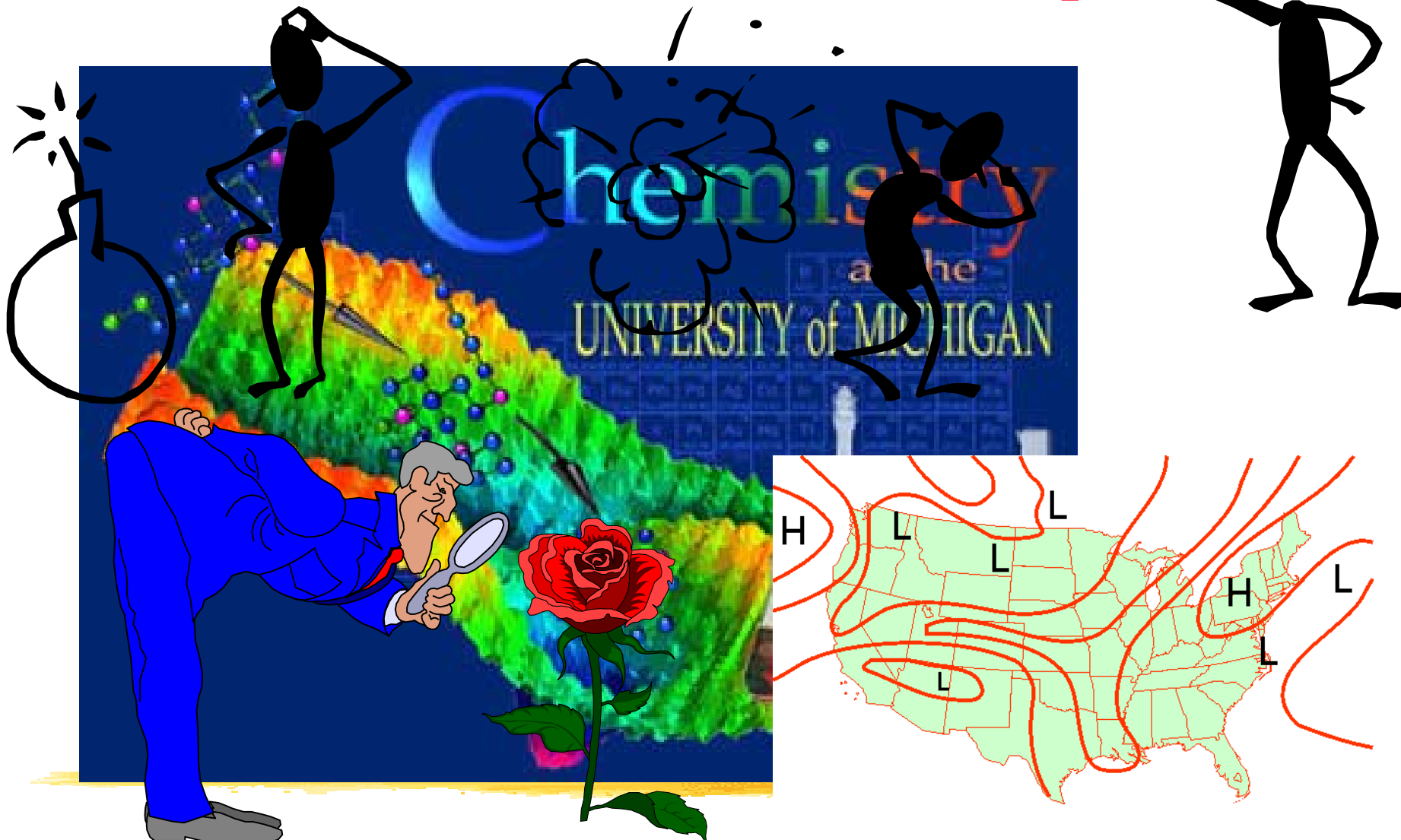
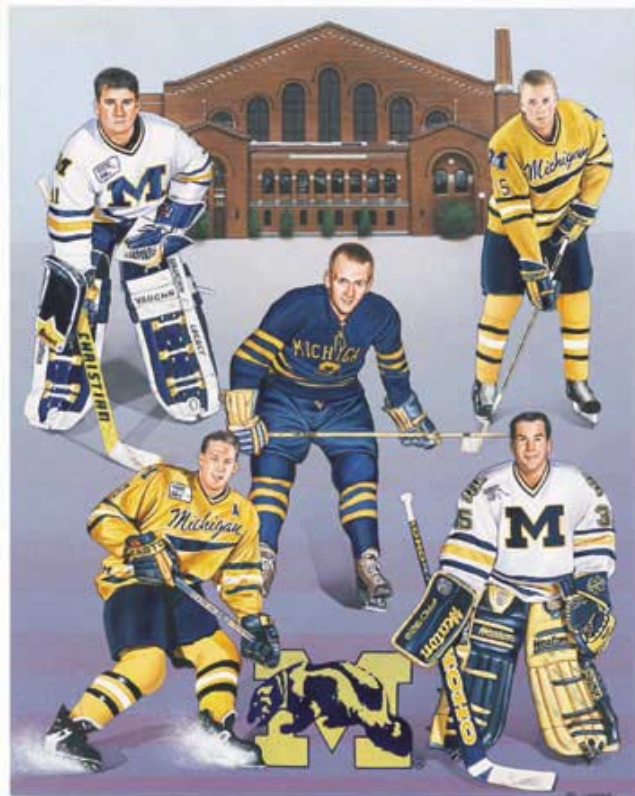


# Chem 260: Chemical Principles

*Nils G. Walter*

**What the h... are Chemical Principles ?**





# Chem 260: Chemical Principles

↓ better:

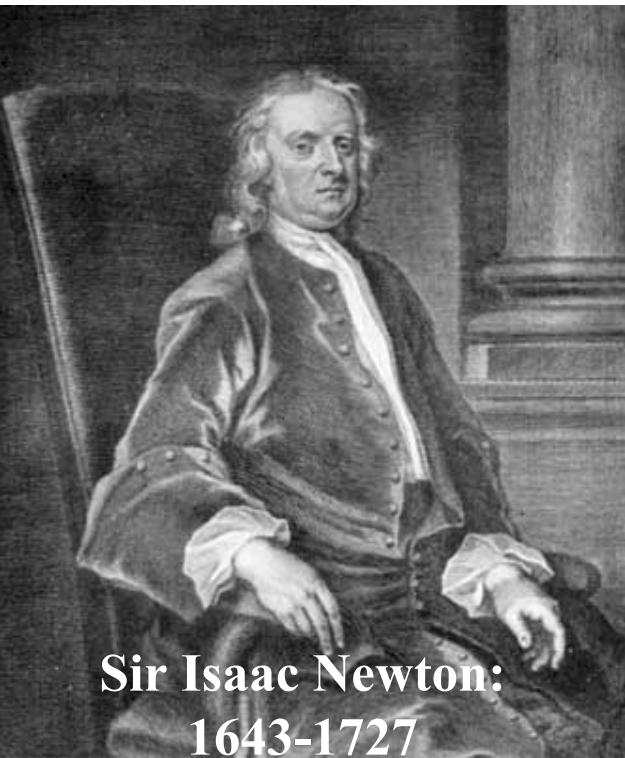
## The Fundamental Physical Principles that Underlie Modern Chemistry and its Applications



# Physics at the End of the 19th Century

**Everything's clear!**

Newton's ideas (published in his *Principia* in 1687) still dominate the physical sciences:



Sir Isaac Newton:  
1643-1727

1. “All matter attracts all other matter with a force proportional to the product of their masses and inversely proportional to the square of their distance” (apples, planets)
2. A particle travels in a trajectory with precise position and momentum
3. Any type of motion can be excited to a state of arbitrary energy
4. Waves and particles are distinct concepts

**Kinetic energy:**  $E_K = \frac{1}{2}mv^2$     **Potential energy:**  $E_P = mgh$

**Force:**  $F = ma$     **Work:**  $w = - \int F(s)ds$     Nils Walter: Chem 260



# The **Dramatic** Failure of Classical Physics

## Case 1: Black-Body Radiation

### Electromagnetic radiation

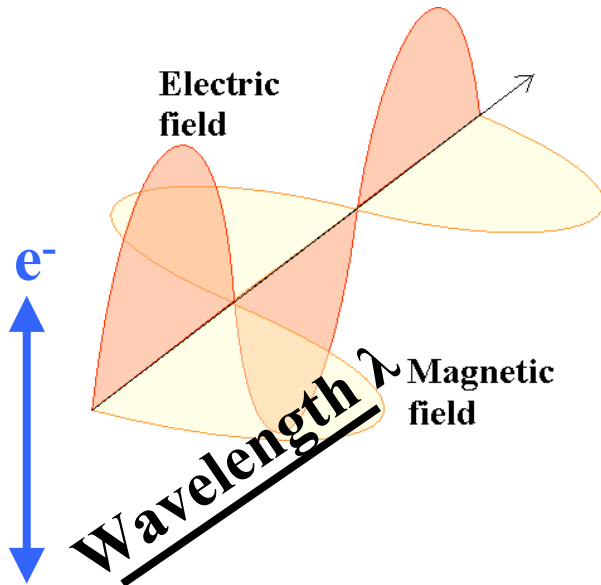


Table 2 Colour, frequency, and wavelength of light<sup>†</sup>

	Frequency/ ( $10^{14}$ Hz)	Wavelength/ nm	Energy of photon/( $10^{-19}$ J)
X-rays and $\gamma$ -rays	$10^3$ and above	3 and below	660 and above
Ultraviolet	10	300	6.6
Visible light			
Violet	7.1	420	4.7
Blue	6.4	470	4.2
Green	5.7	530	3.7
Yellow	5.2	580	3.4
Orange	4.8	620	3.2
Red	4.3	700	2.8
Infrared	3.0	1000	1.9
Microwaves and radiowaves	$3 \times 10^{-11}$ Hz and below	$3 \times 10^6$ and above	$2.0 \times 10^{-22}$ J and below

<sup>†</sup> The values given are approximate but typical.

Frequency  $\nu$

$$\lambda = \frac{c}{\nu} = \frac{1}{\tilde{\nu}}$$

Speed of light

Wavenumber

What has this to do with a black body  
(capable to emit  
and absorb all  $\nu$ )?

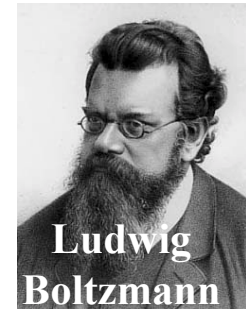
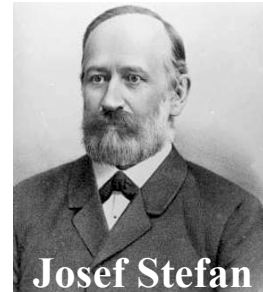
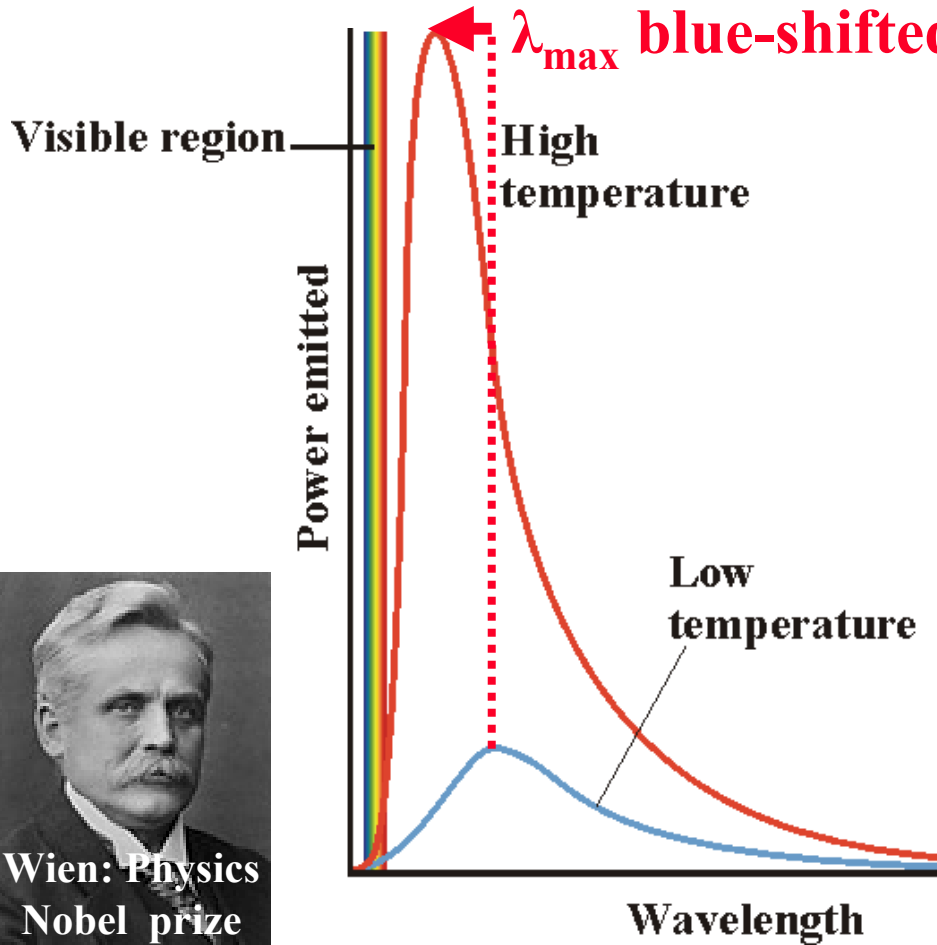


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# Properties of Black-Body Radiation

Power density of a black body:

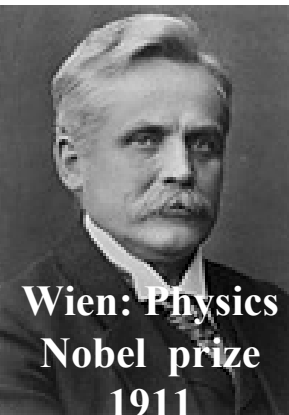


Emittance = total power emitted  
 $\Rightarrow$  Stefan-Boltzmann law:

$$M = aT^4; a = 56.7 \text{ nW m}^{-2} \text{ K}^{-4}$$

$\Rightarrow$  The higher the temperature of a lamp wire, the more power will be emitted as light!

$\Rightarrow$  Halogen lamps!



Wien's displacement law:  $T\lambda_{\text{max}} = 2.9 \text{ mm K}$

$\Rightarrow$  Surface temperature of the sun w/  $\lambda_{\text{max}} \approx 490 \text{ nm!}$