

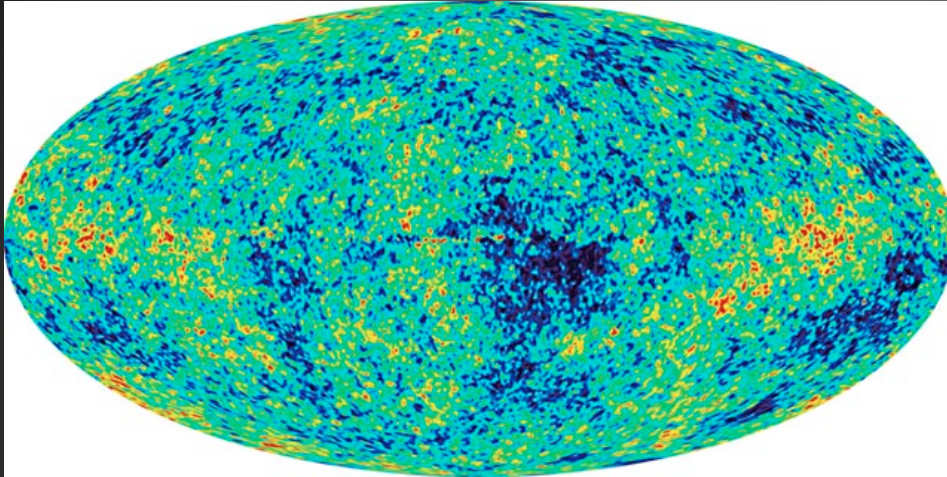
Fabrication, Assembly, and First  
Measurements of a New Facility to  
Characterize Optical Elements for  
Observing the Cosmic Microwave  
Background

Kurt Flesch

University of Wisconsin-Eau Claire

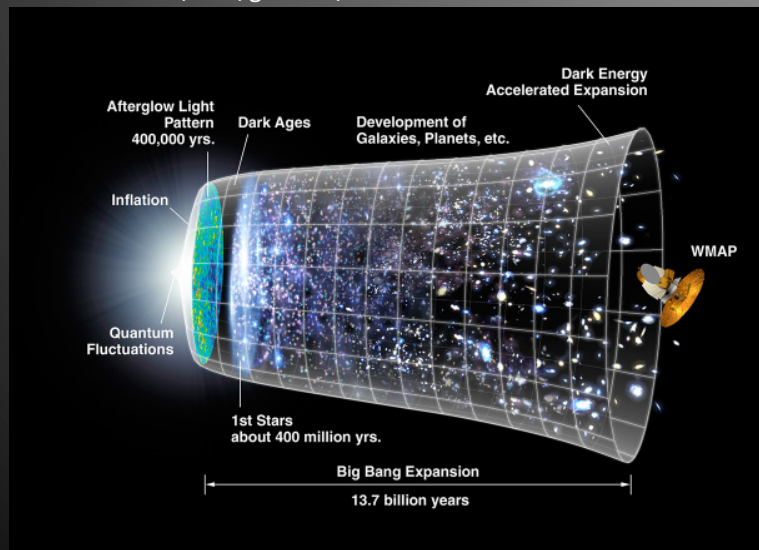
Advisor: Dr. Jeff McMahan

# The Cosmic Microwave Background



<http://www.astr.ua.edu/keel/galaxies/cmbr.html>

- Radiation left from Big Bang
- Roughly 2.7K in all directions
- Small temperature fluctuations

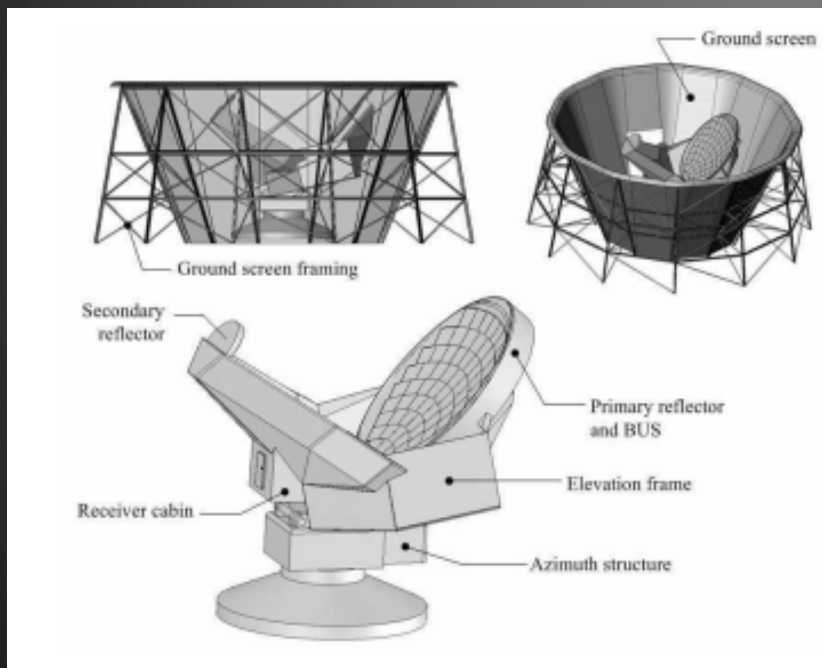


[http://www.nasa.gov/vision/universe/starsgalaxies/wmap\\_pol.html](http://www.nasa.gov/vision/universe/starsgalaxies/wmap_pol.html)

# Goals of ACTPol

(Atacama Cosmology Telescope)

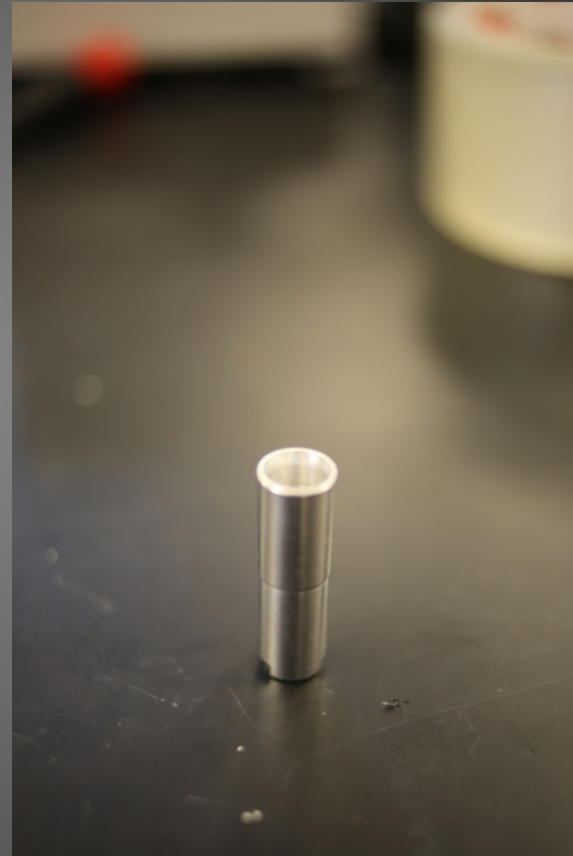
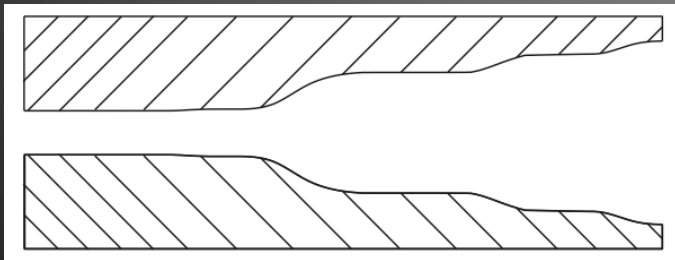
- Orientation of polarization
- Multiple frequency pixels
- Learn more about expansion of the universe, neutrino masses, and the nature of dark matter and dark energy



<http://arxiv.org/pdf/astro-ph/0701020.pdf>

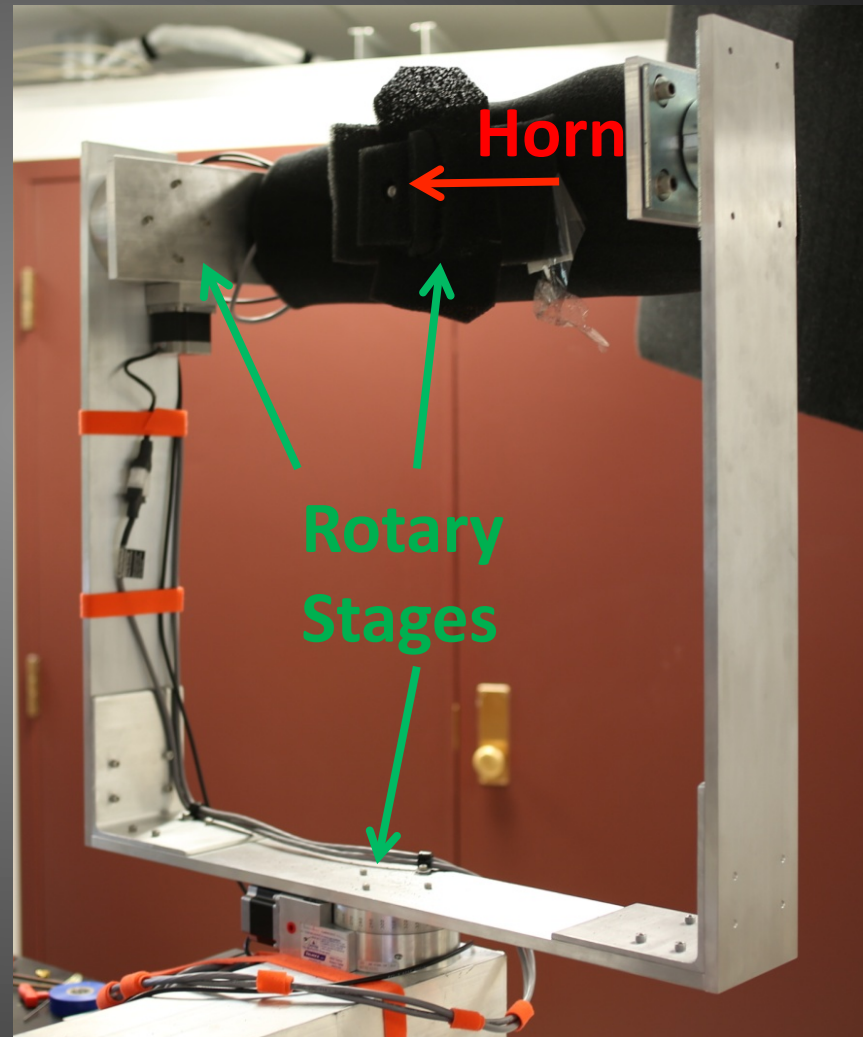
# Beam Mapping

- Diffraction Pattern
- Polarization Coupling
- Interpretation of data from the telescope



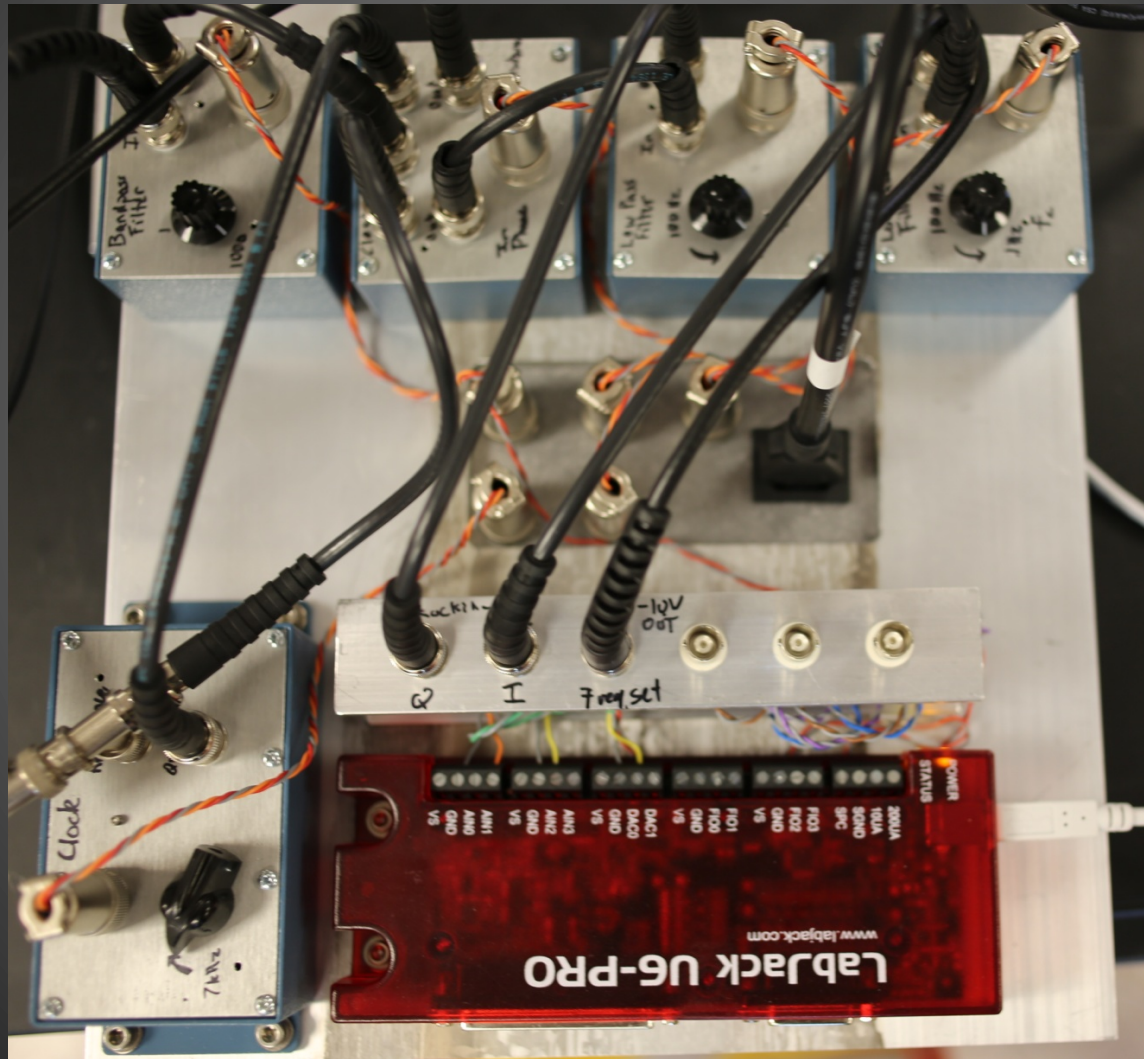
# Horn Mount

- Large design
- Microwave absorbing foam
- Three axes of rotation
- Front of horn located on axes of rotation

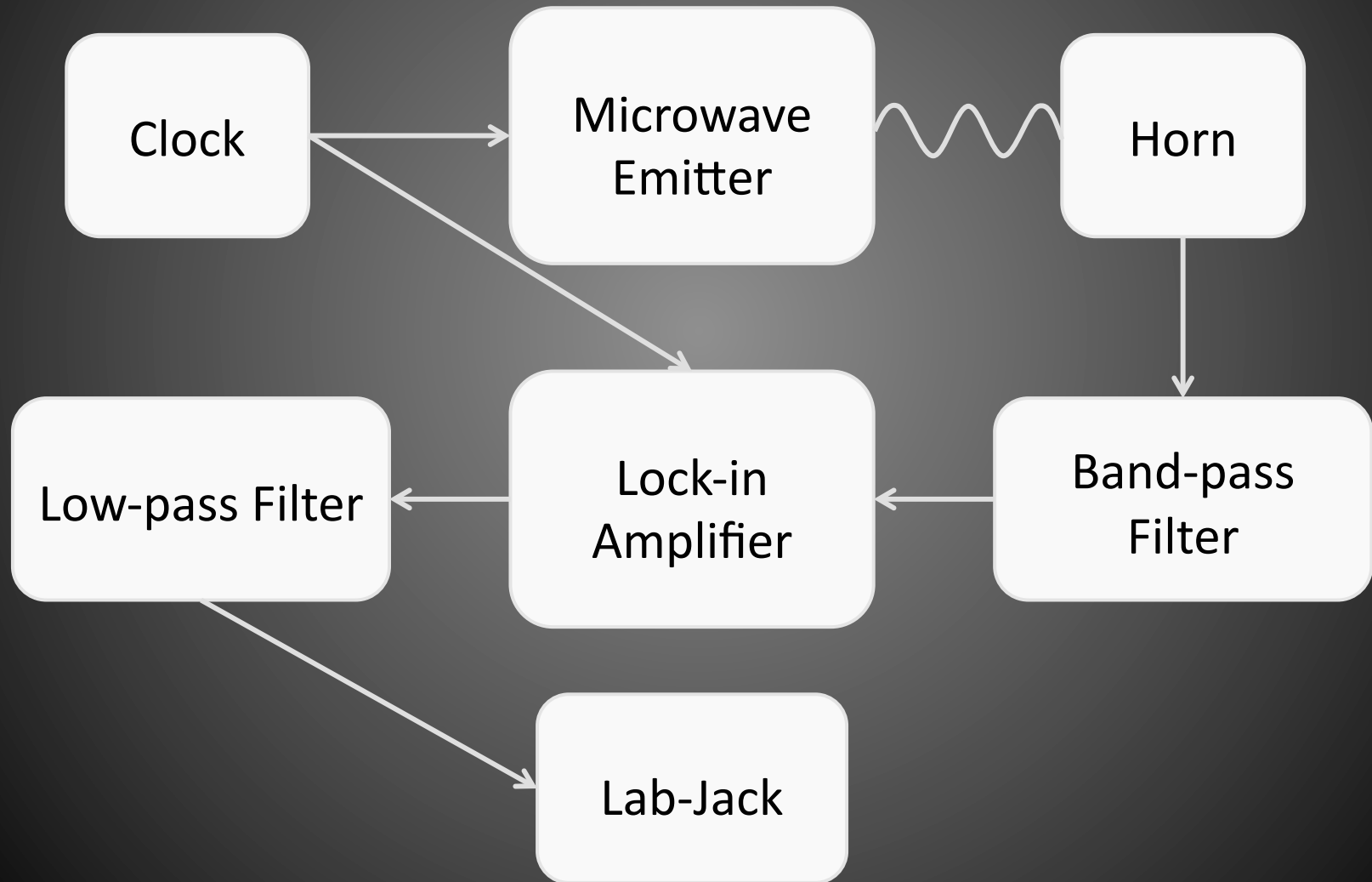




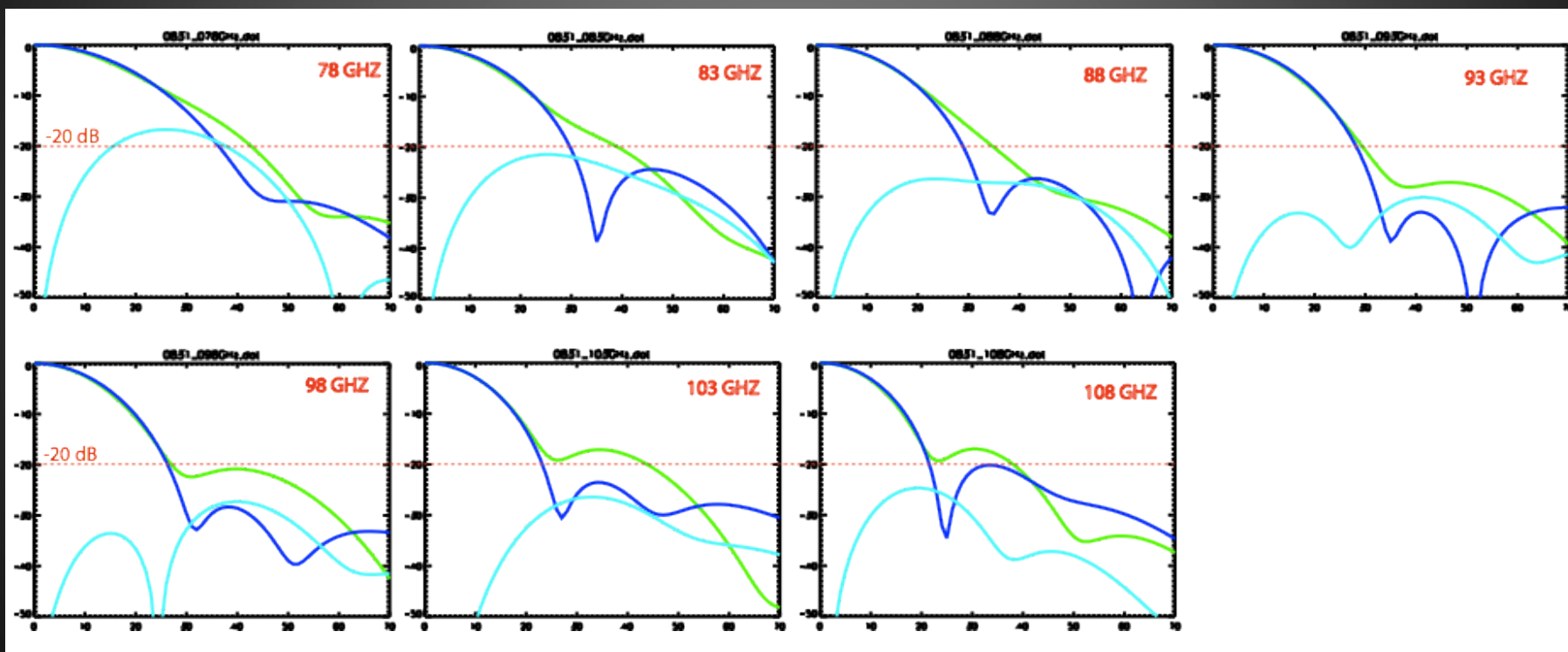
# Circuit Plate



# Circuit Setup

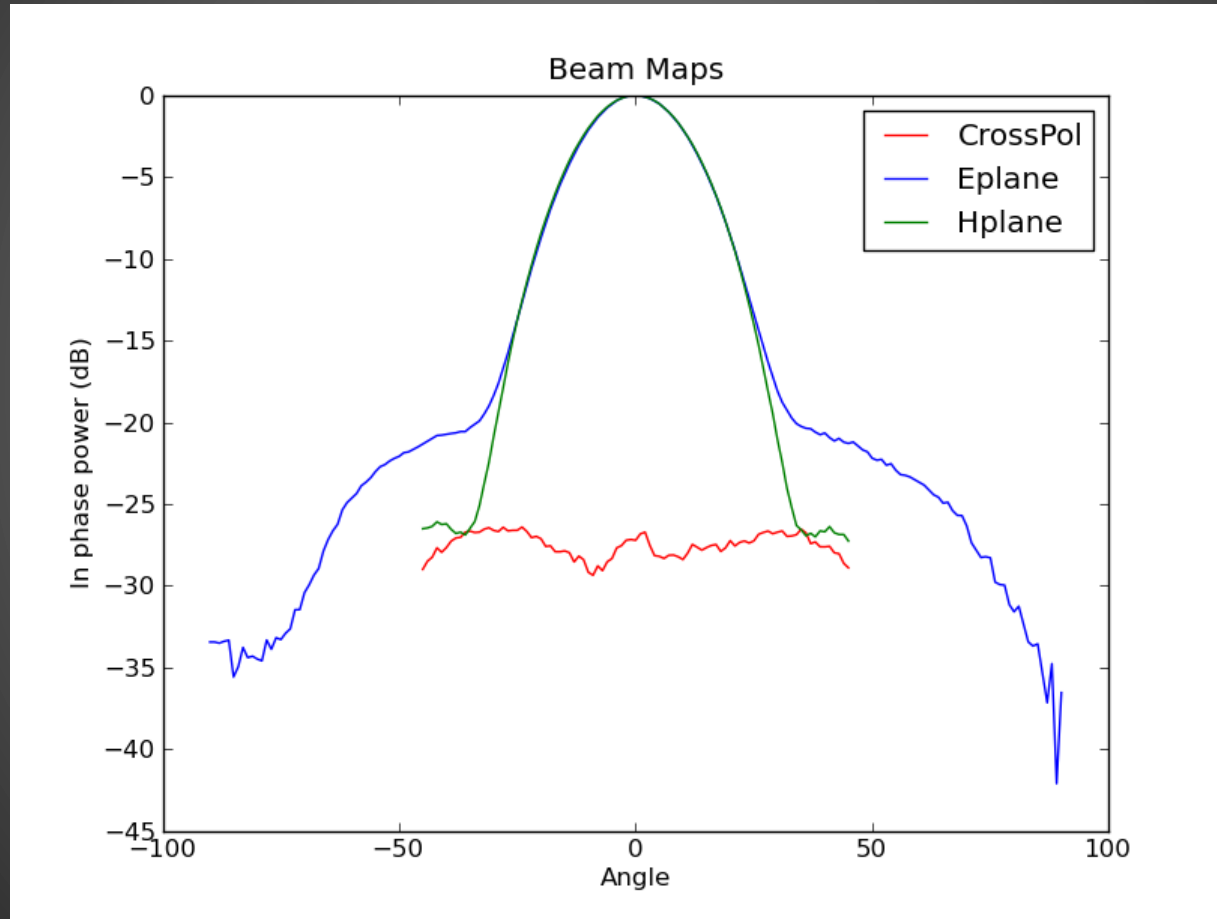


# Modeled Data

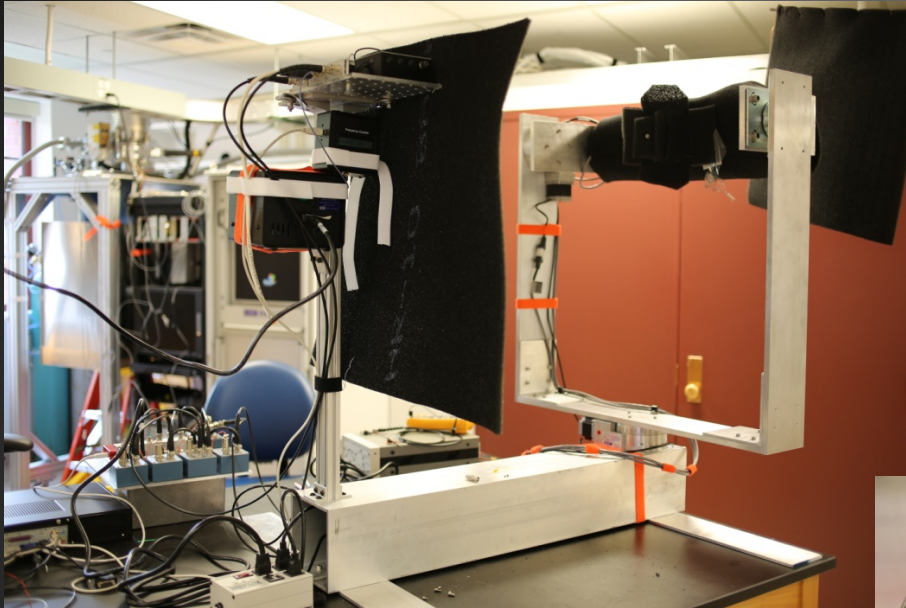




# Results

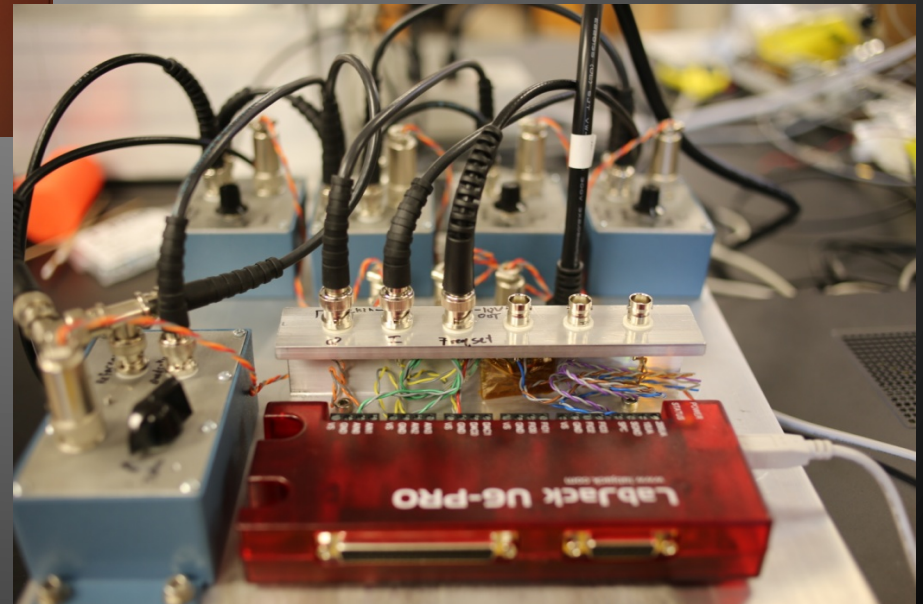


# Future Work



- Can be kept in the lab and used to test new horns

- Continue to reduce reflections and improve computer coding



# Acknowledgements

- Dr. Jeff McMahon
- Charles Munson
- University of Michigan Physics
- NSF



Questions?