

Spot-o-Matic

Testing Intrapixel Variation in the SNAP Detectors

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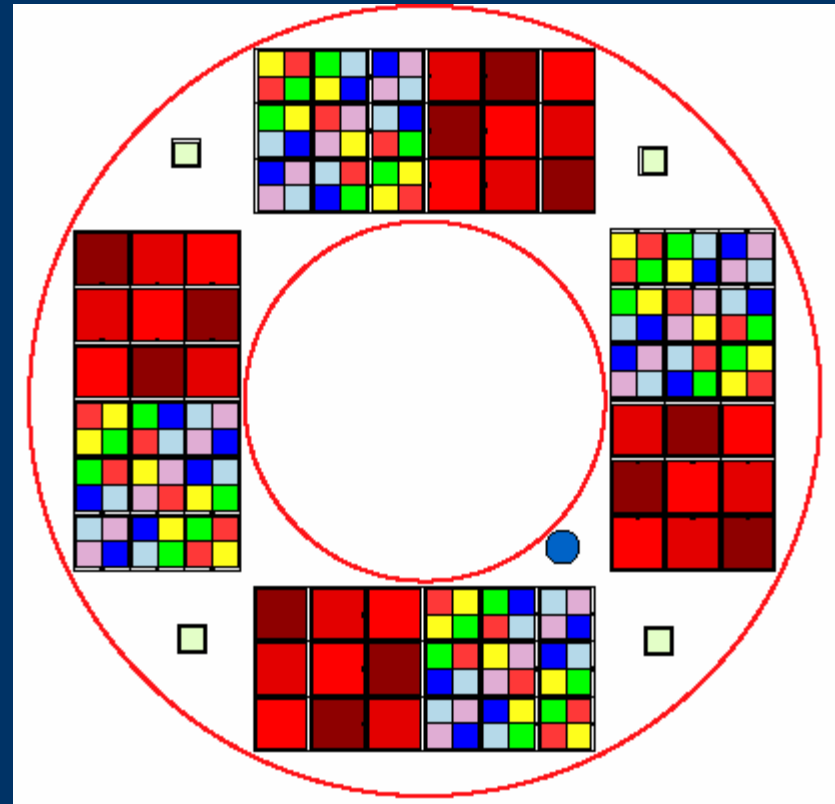
REU 2004

Spot-o-matic

SNAP Detector Design

36 HgCdTe 2k x 2k NIR
detectors with $18\mu\text{m}$
pixels

Allows us to see all
the way to $z=1.7\mu\text{m}$



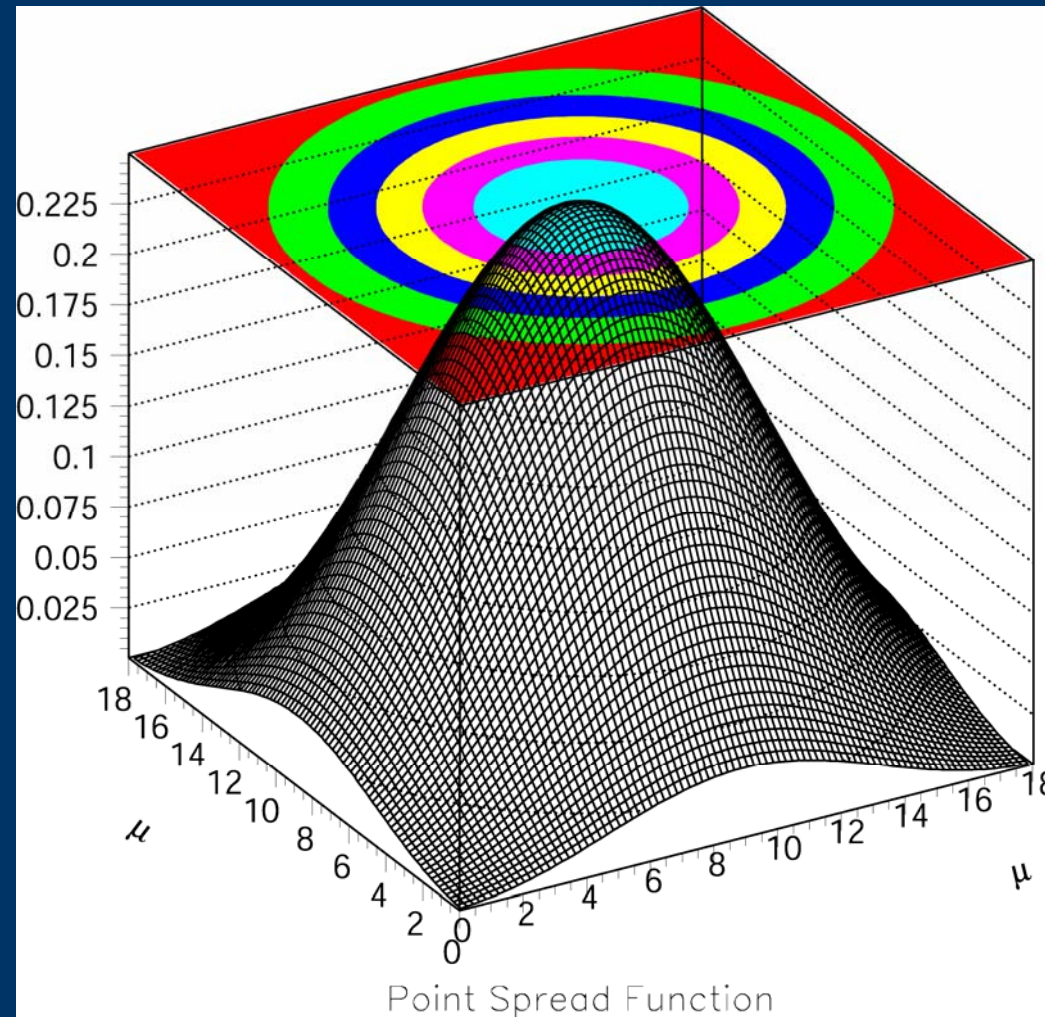
The Problem of Undersampling

SNAP uses $18\mu\text{m}$ pixels

We need greater resolution than this, or else any structure under $18\mu\text{m}$ is lost!

An image of a galaxy which is $40\mu\text{m}$ wide will not look like a galaxy, it'll look like 4 lit up pixels!

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Airy disk for $\lambda = 1 \mu\text{m}$

There are ways around this problem!

- Dithering
- Dithering requires accurate knowledge of the pixel response function (QE vs. position), which we do not know

Enter the Spot-o-Matic!

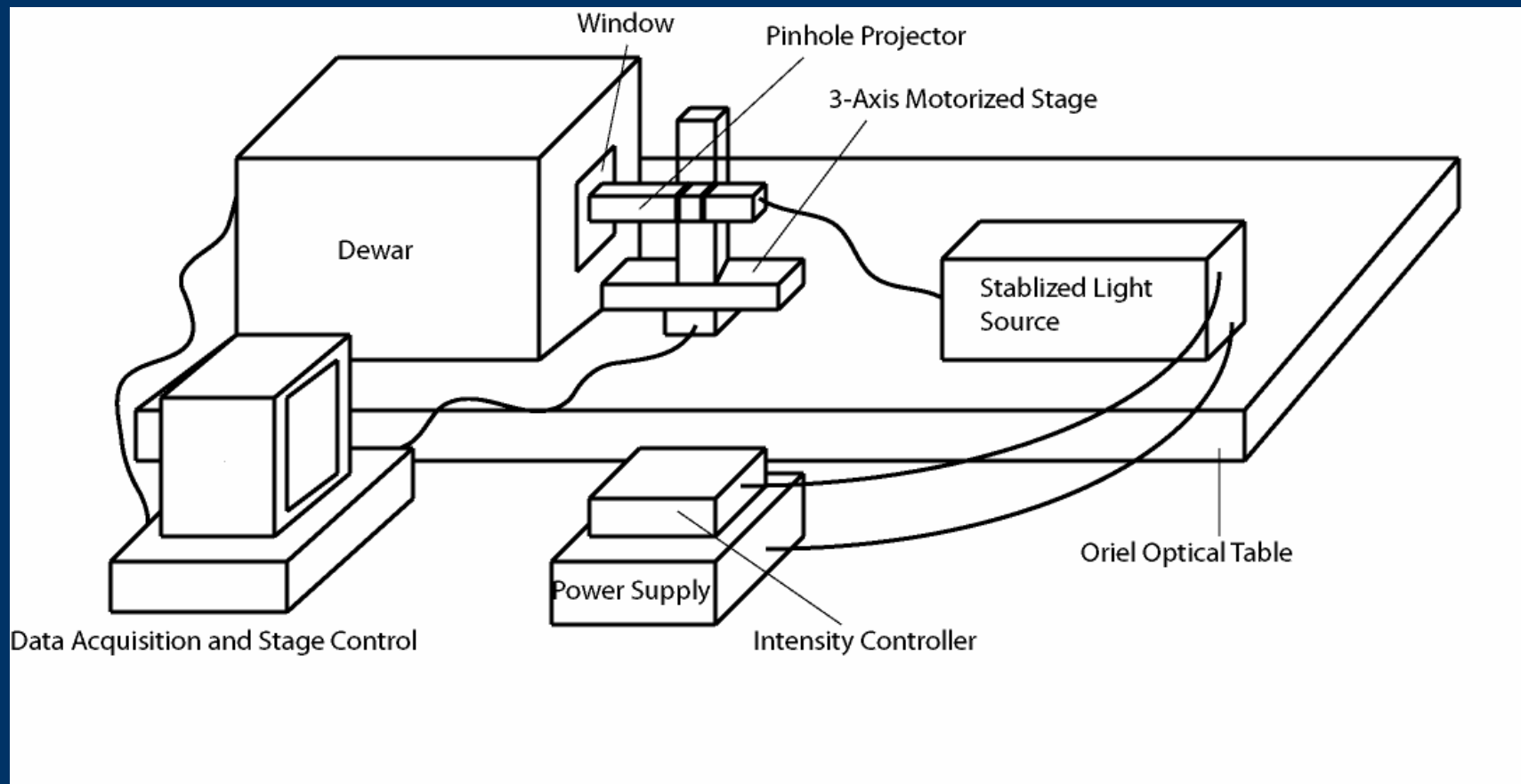
The pixel response function can be extracted by shining a very small spot of NIR light onto the detector, and measuring the readout at different points on the pixel itself!

Testing Setup



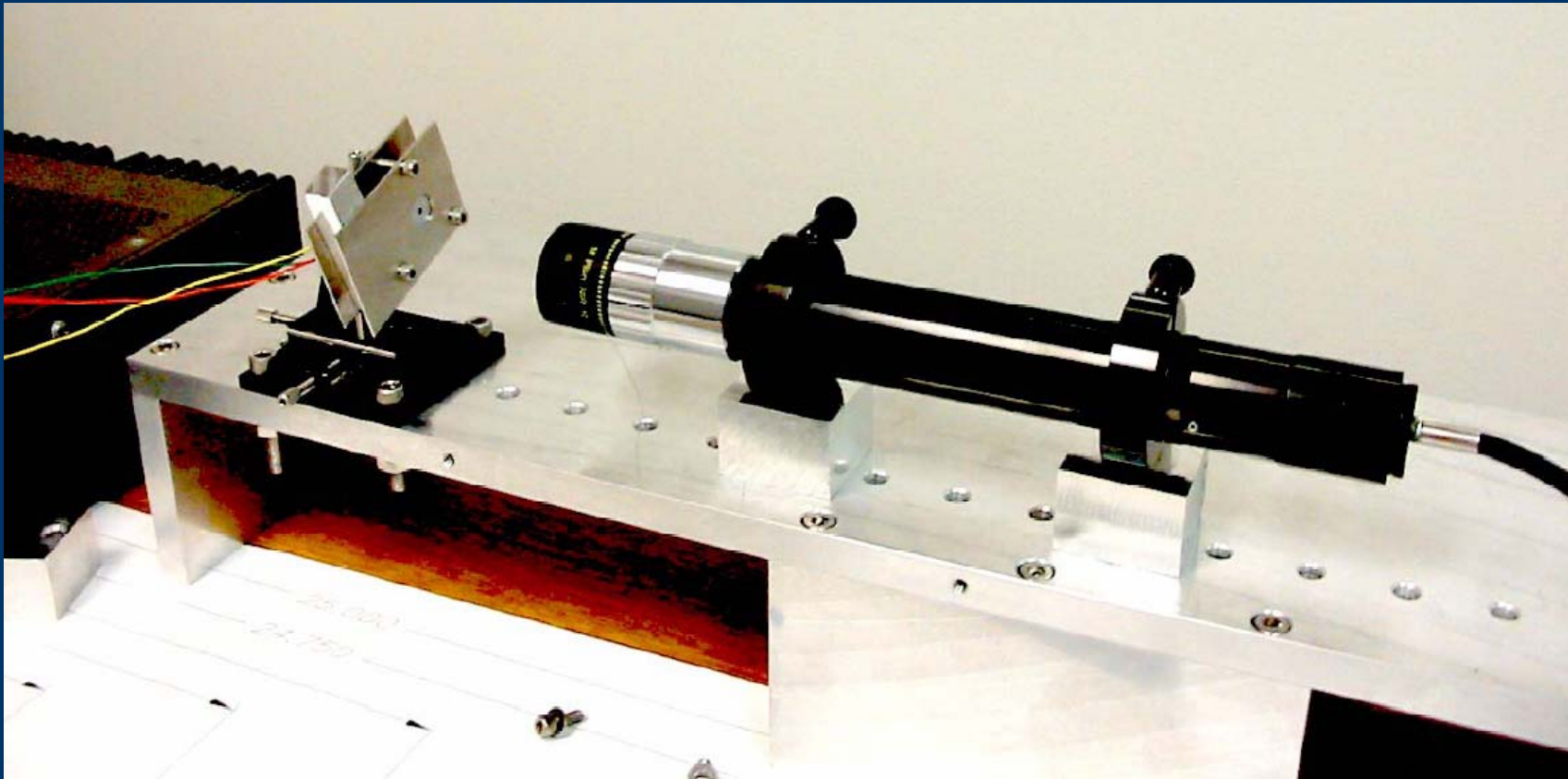
Devices are mounted in a dewar and kept at 140K to reduce dark current and simulate deep space conditions

Testing Setup



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Progress as of REU 2003:



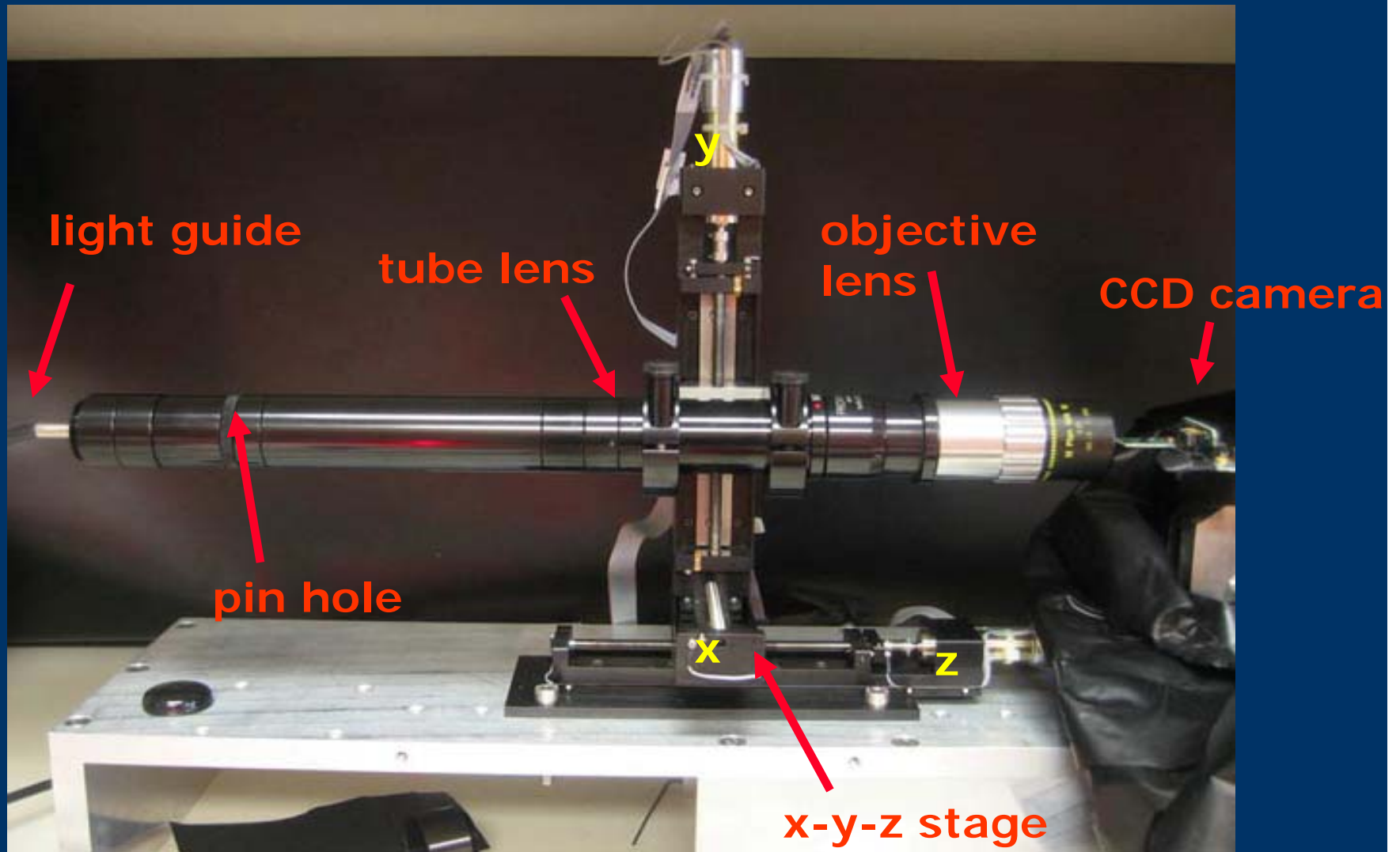
Spot-o-Matic

SNAP

SuperNova
Acceleration
Probe

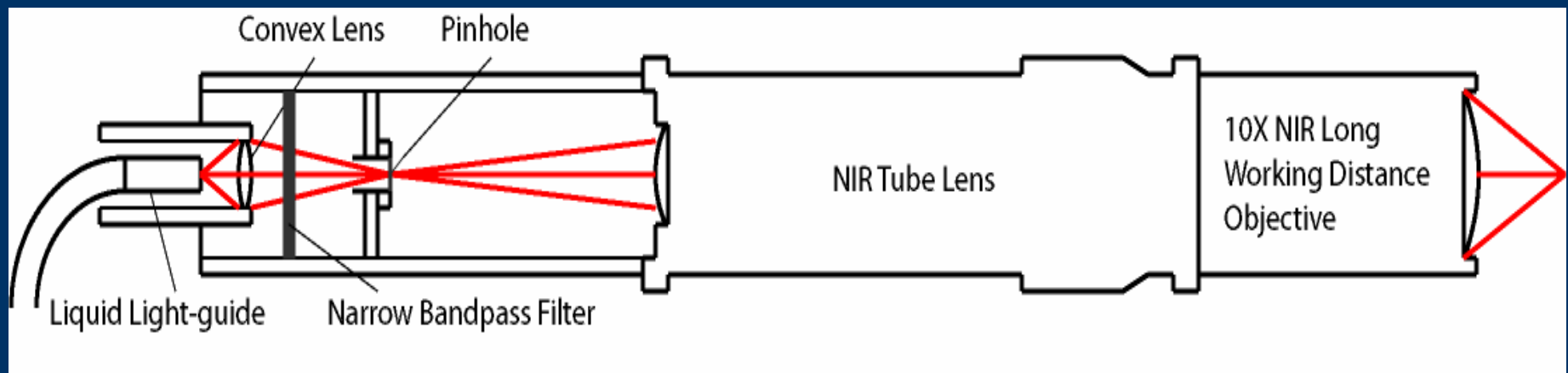
at the University of Michigan

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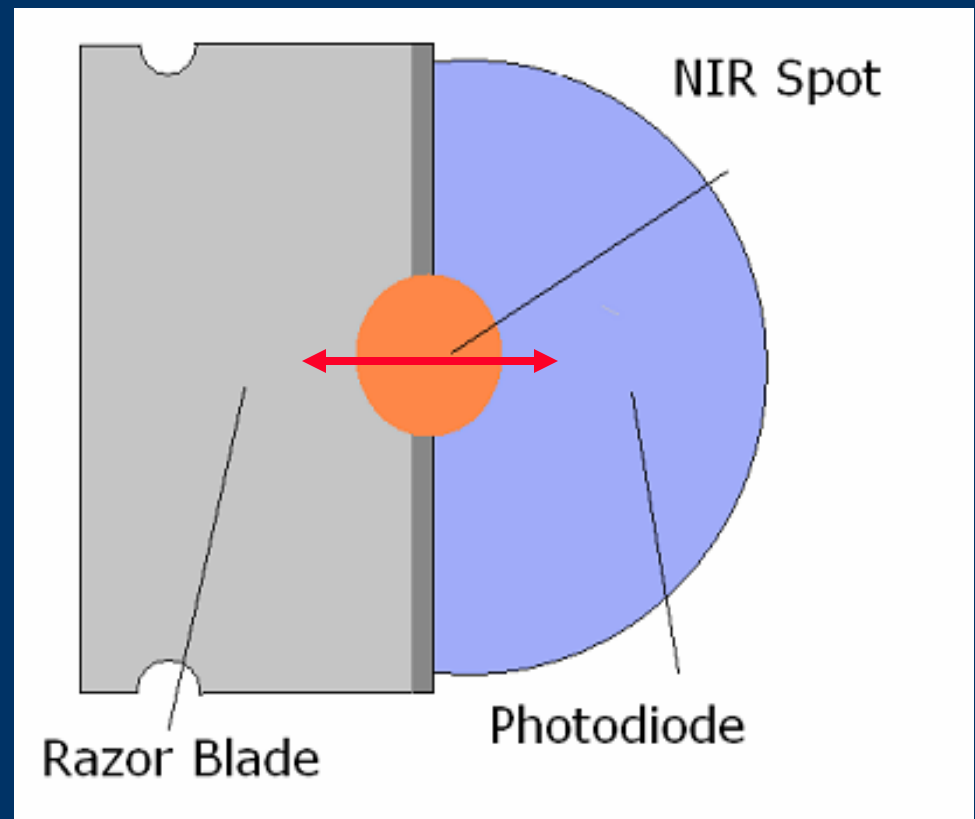
How it works



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Accurately determining the PSF

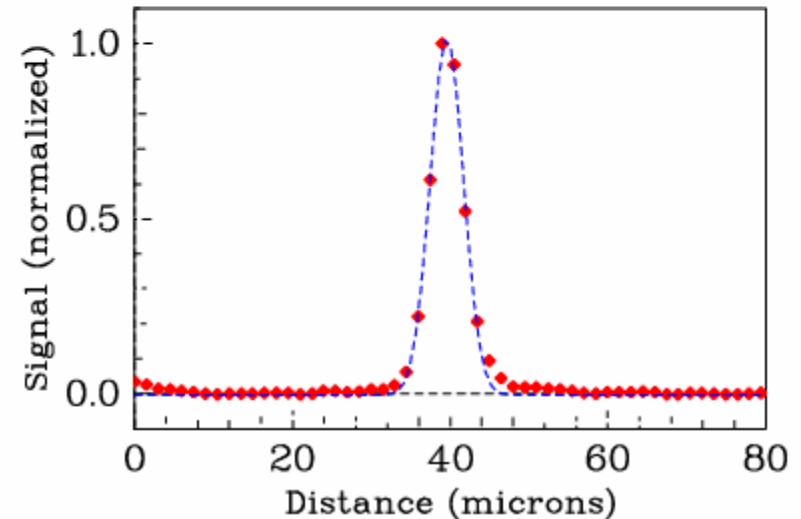
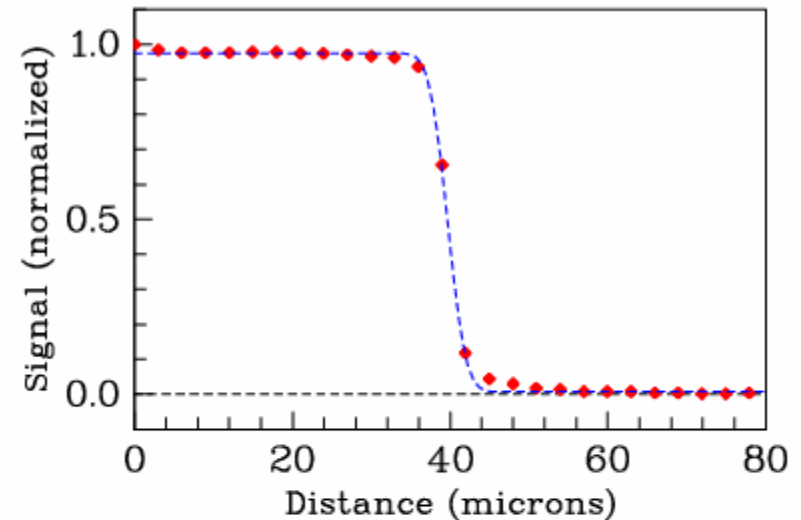
In order to determine the response of a pixel, we need to know the exact size of the spot that we shine on the detector



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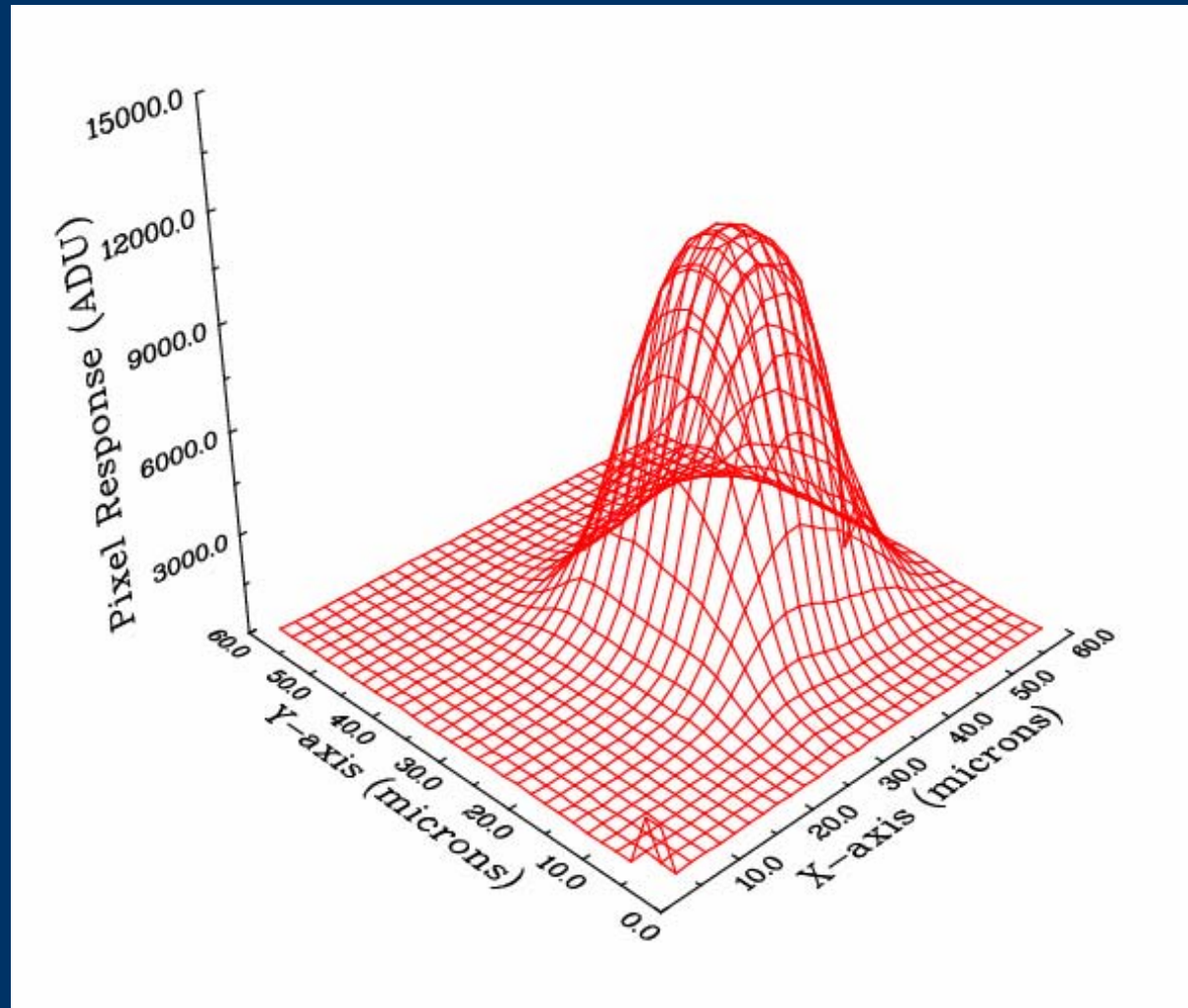
At $\lambda = 1550\text{nm}$, we have a spot profile with $\sigma = 2.10 \mu\text{m}$

This is definitely small enough to probe a $18 \mu\text{m}$ pixel



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Pixel Scan



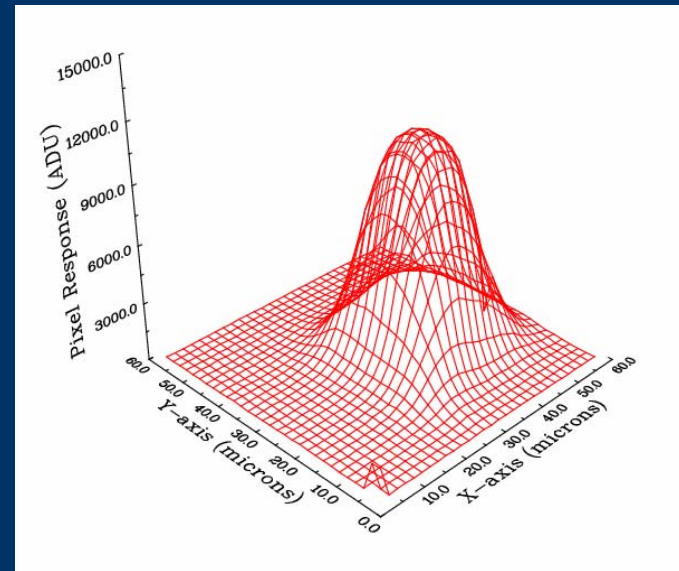
What do we notice about the pixel?

- Pixel response is quite smooth and uniform over its area (no significant dip in QE)
- The response of the pixel extends beyond the physical size of the pixel itself!

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What does this mean?

Lateral charge diffusion



This doesn't mean we can't dither to achieve precision photometry, as long as we know exactly how much charge diffusion we have

Deconvolution of PRF

The results we obtained are not actually the PRF, but rather a convolution of the PRF and the PSF

Deconvolution is necessary to determine the true PRF

The future of the Spot-o-Matic

- UM now has the only NIR spot projection system with the ability to precisely resolve the PRF of these detectors
- Can be used for a variety of experiments, including simulation of an actual PSF that SNAP will see in space

Special Thanks to:

- Wolfgang Lorenzon
- Michael Borysow
- NSF and the UM REU Program
- The entire SNAP team