

An LBT view of faint Milky Way dwarf galaxies

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Hercules

Canes Venatici I

Leo T

The Large Binocular Telescope Hill et al. (2006)

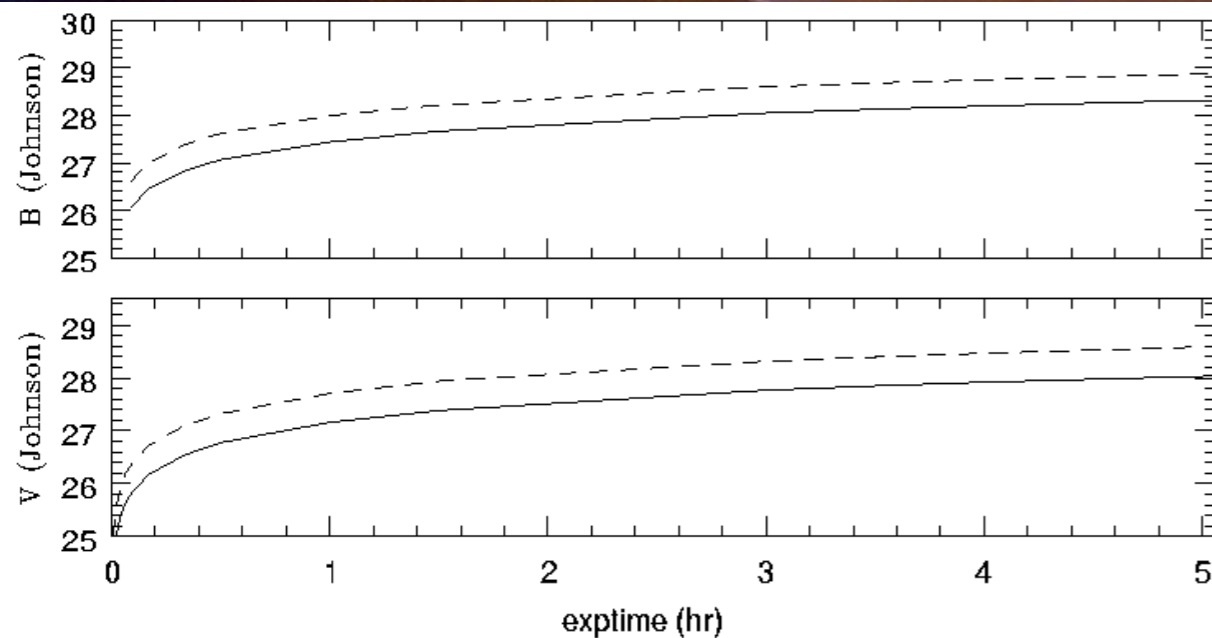
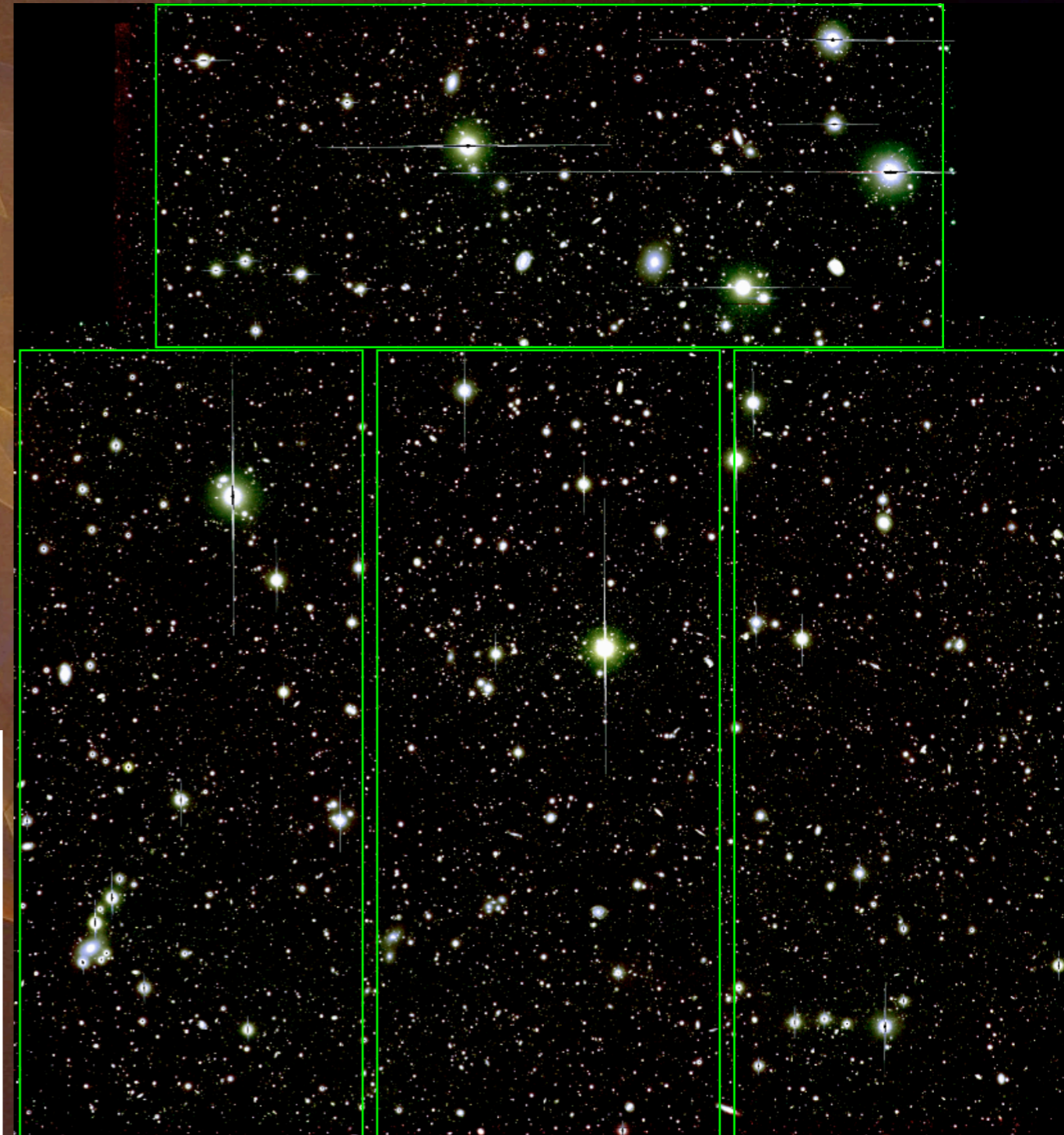
- * Mt Graham Int. Obs., Arizona
- * 2 x 8.4m mirrors on common mount → 11.8 m aperture
- * First light:
 - ★ Mirror 1 in October 2005
 - ★ Mirror 2 in September 2006
- * Currently 'blue' side fitted with Large Binocular Camera



The Large Binocular Camera Blue

Giallongo et al. (2006)

- * Wide-Field Imager
 - ★ Four 2048 x 4608 chips
 - ★ $0.23'' / \text{px} \rightarrow 23' \times 23' \text{ FoV}$
- * 2 LBC to target blue and red simultaneously
- * 5 min. exposure $\rightarrow V \sim 25.5$
- * First instrument on LBT



Hercules

Coleman et al. (2007)

* Belokurov et al. 06:

★ $M_V = -6.0$

★ $D = 140$ kpc

* Simon & Geha 07:

★ $v_{r, \text{gsr}} = 145$ km/s

★ $\sigma = 5.1$ km/s

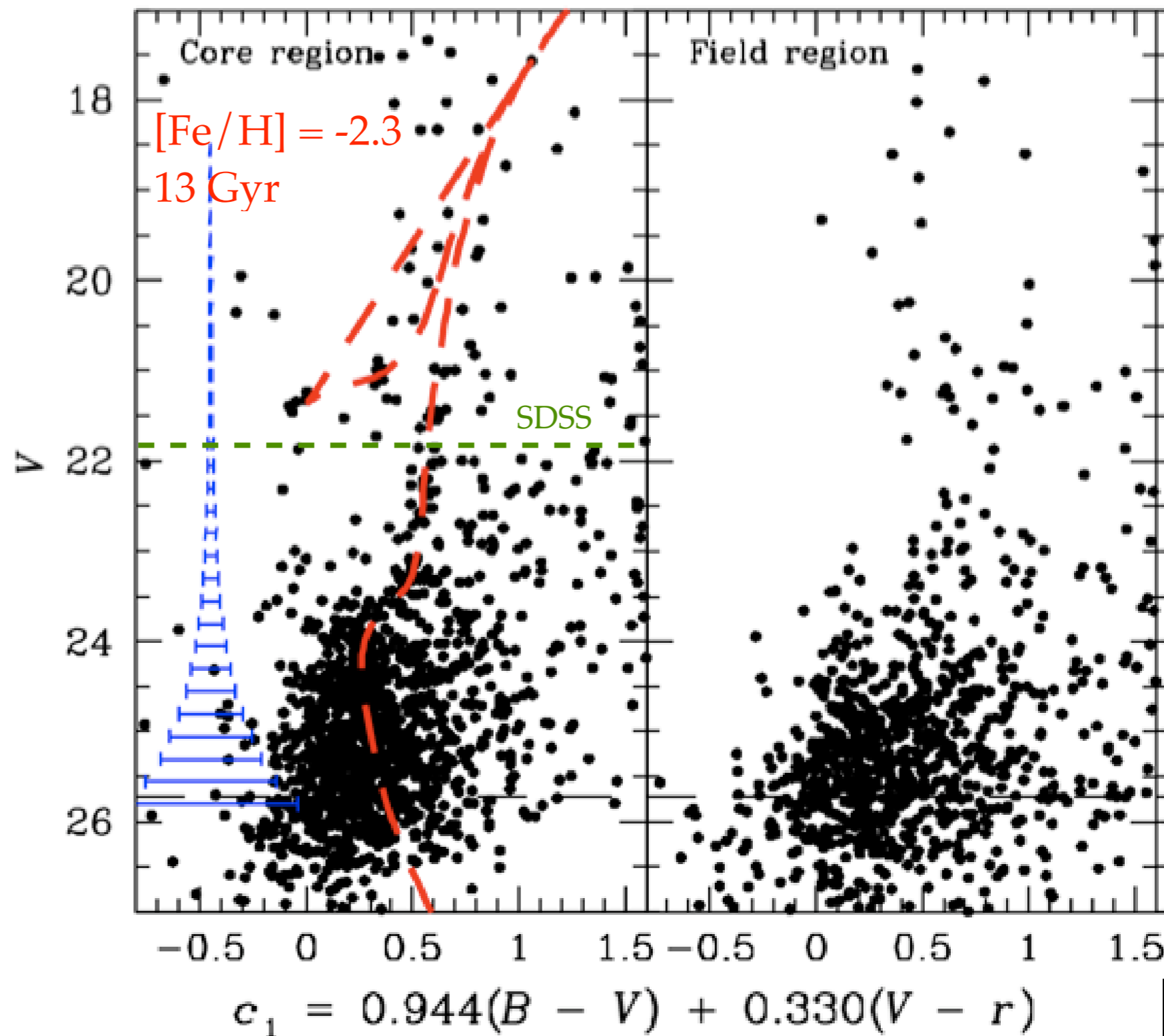
★ $[\text{Fe}/\text{H}] = -2.3$

* LBC:

★ $30' B, 20' V, 20' r$

★ 50% comp. at $V \sim 25.5$

★ seeing $0.8 - 1.1''$

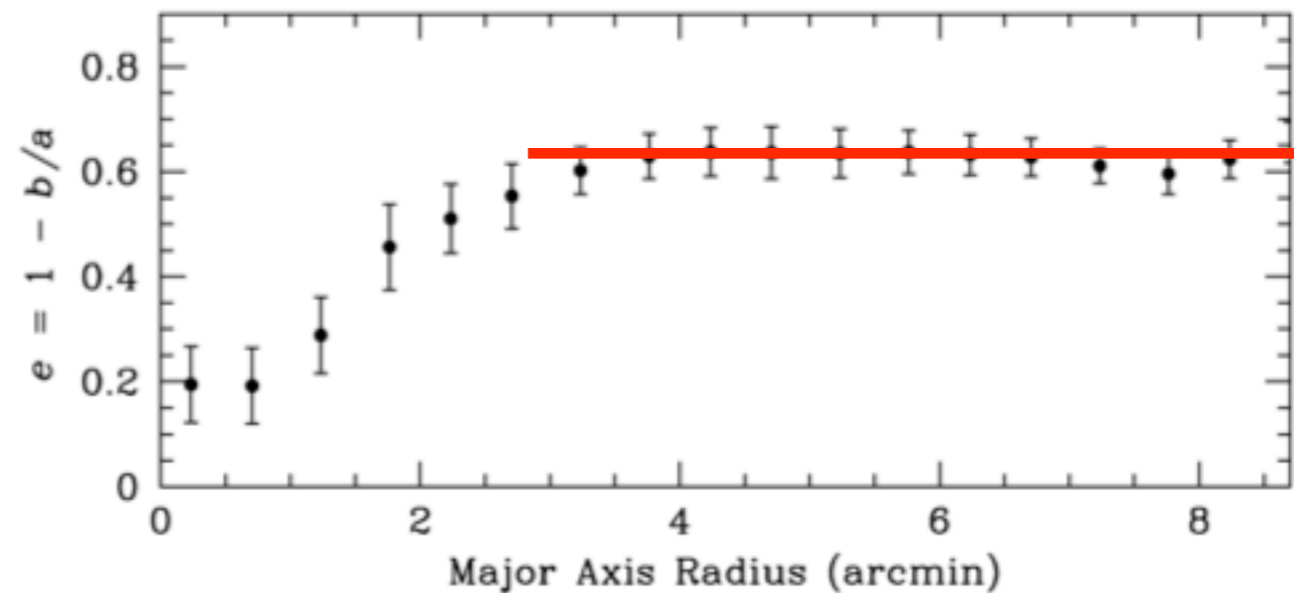
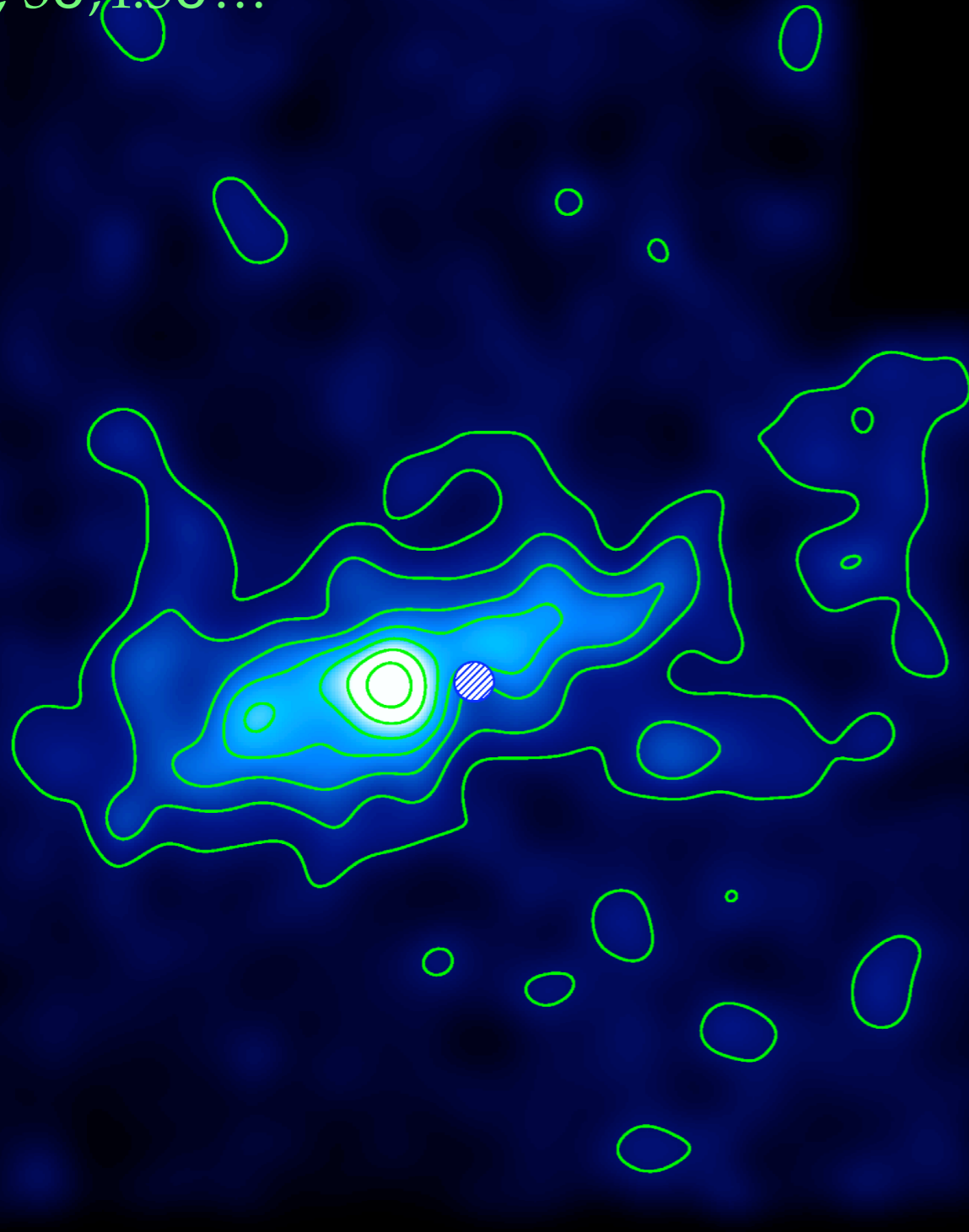


Hercules

Coleman et al. (2007)

✧ *Very flat*, axis ratio $\sim 3:1$

1.5 σ , 3 σ , 4.5 σ ...



✧ CMD fitting:

★ 132 ± 12 kpc

★ $[Fe/H] = -2.1 \pm 0.2$

★ 13 ± 3 Gyr

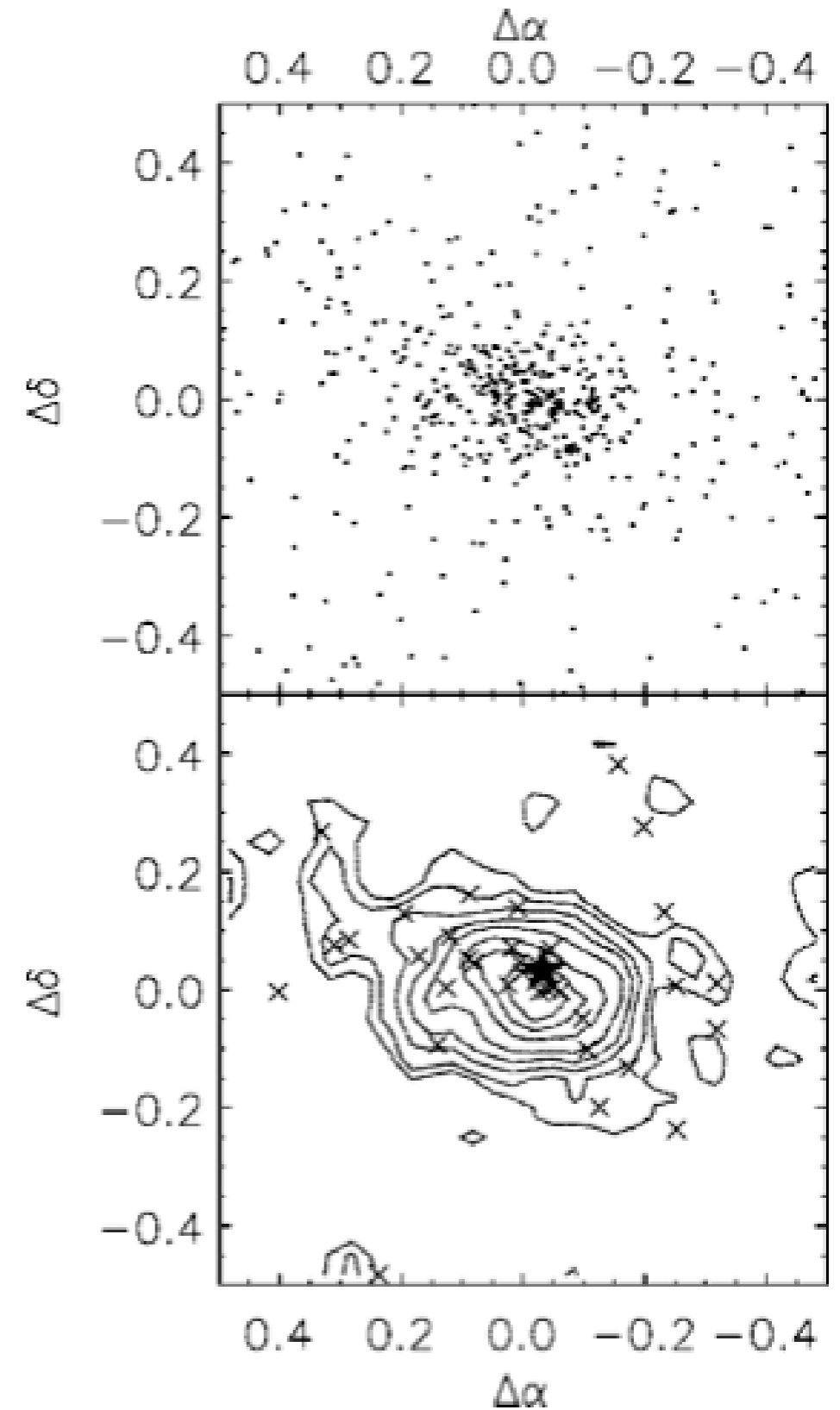
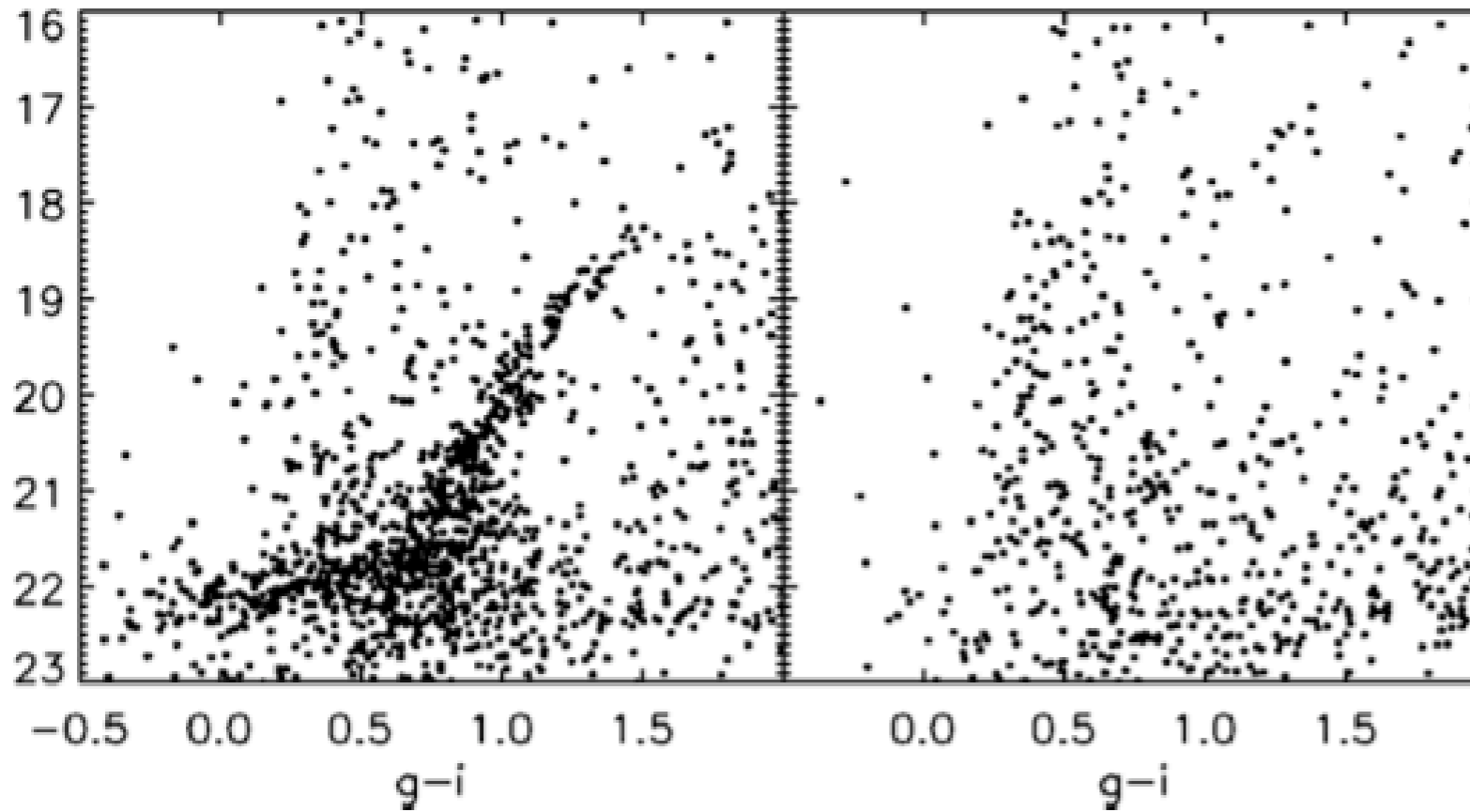
★ $r_h = 168 \pm 11$ pc

$M_V = -7.9$

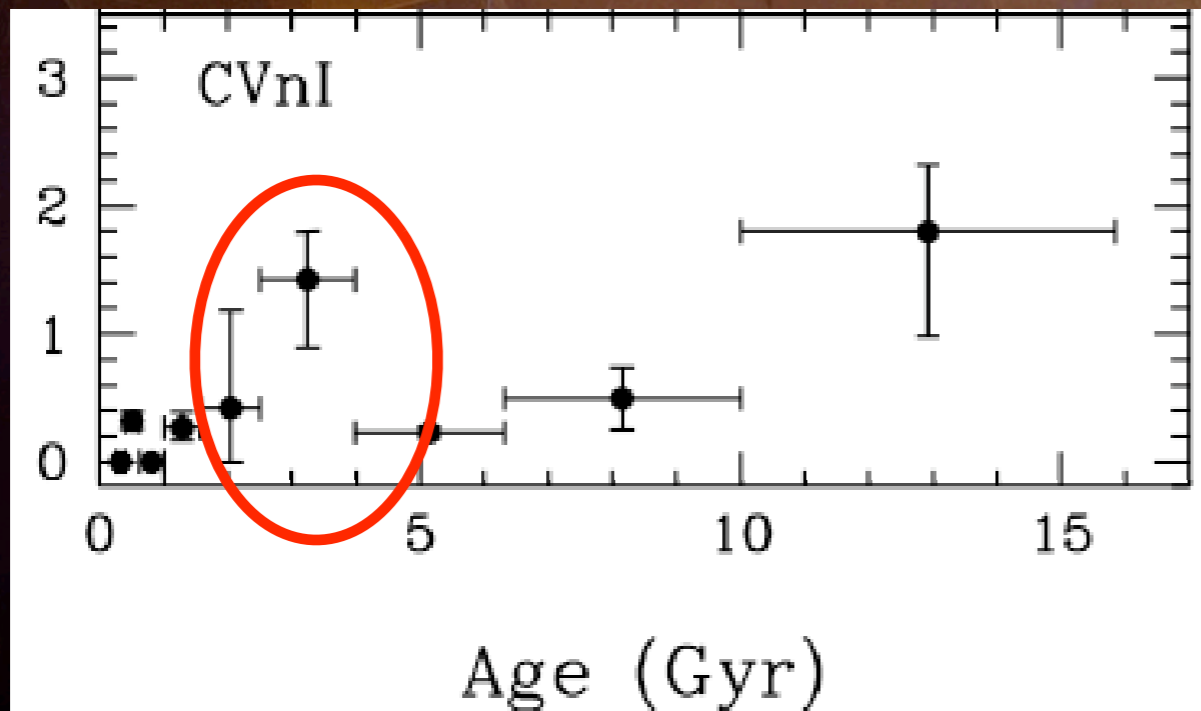
$D = 220$ kpc

Canes Venatici I (SDSS)

Zucker et al. (2006)



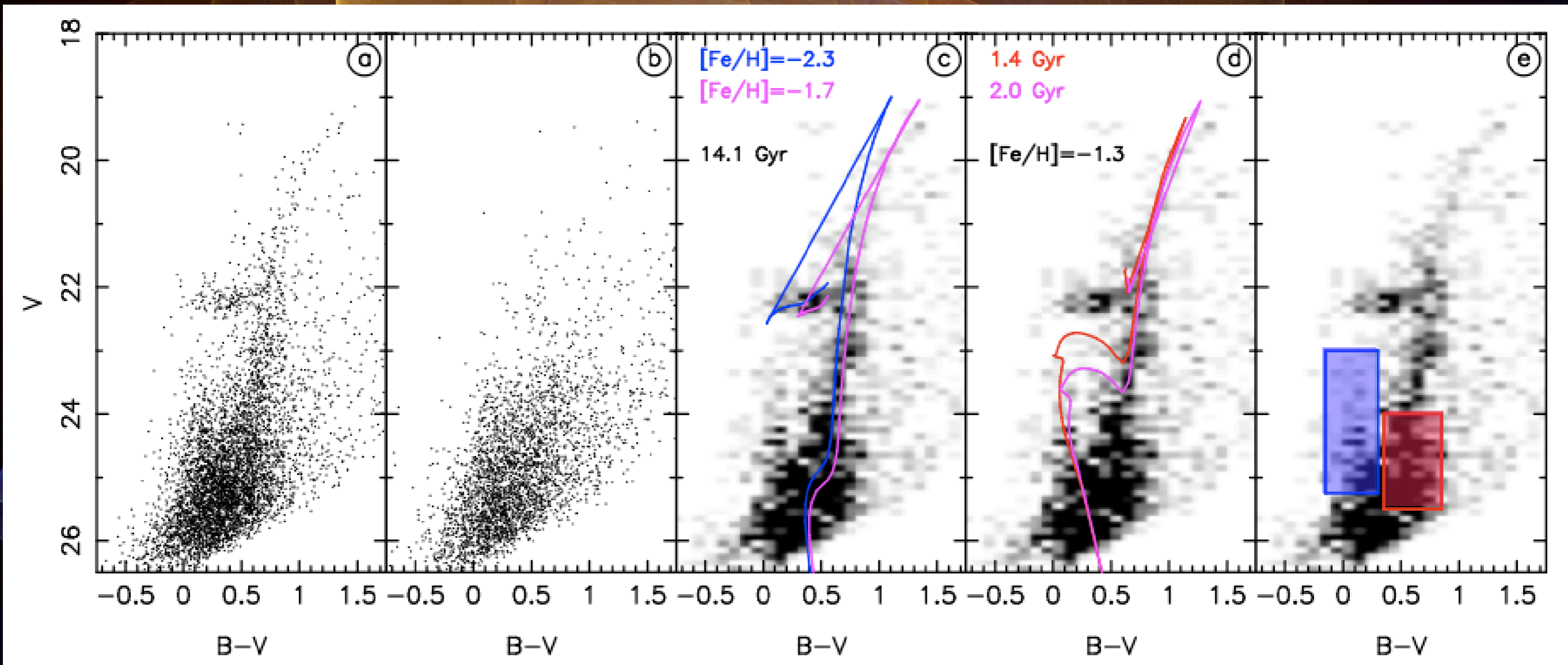
de Jong et al. (2007a)

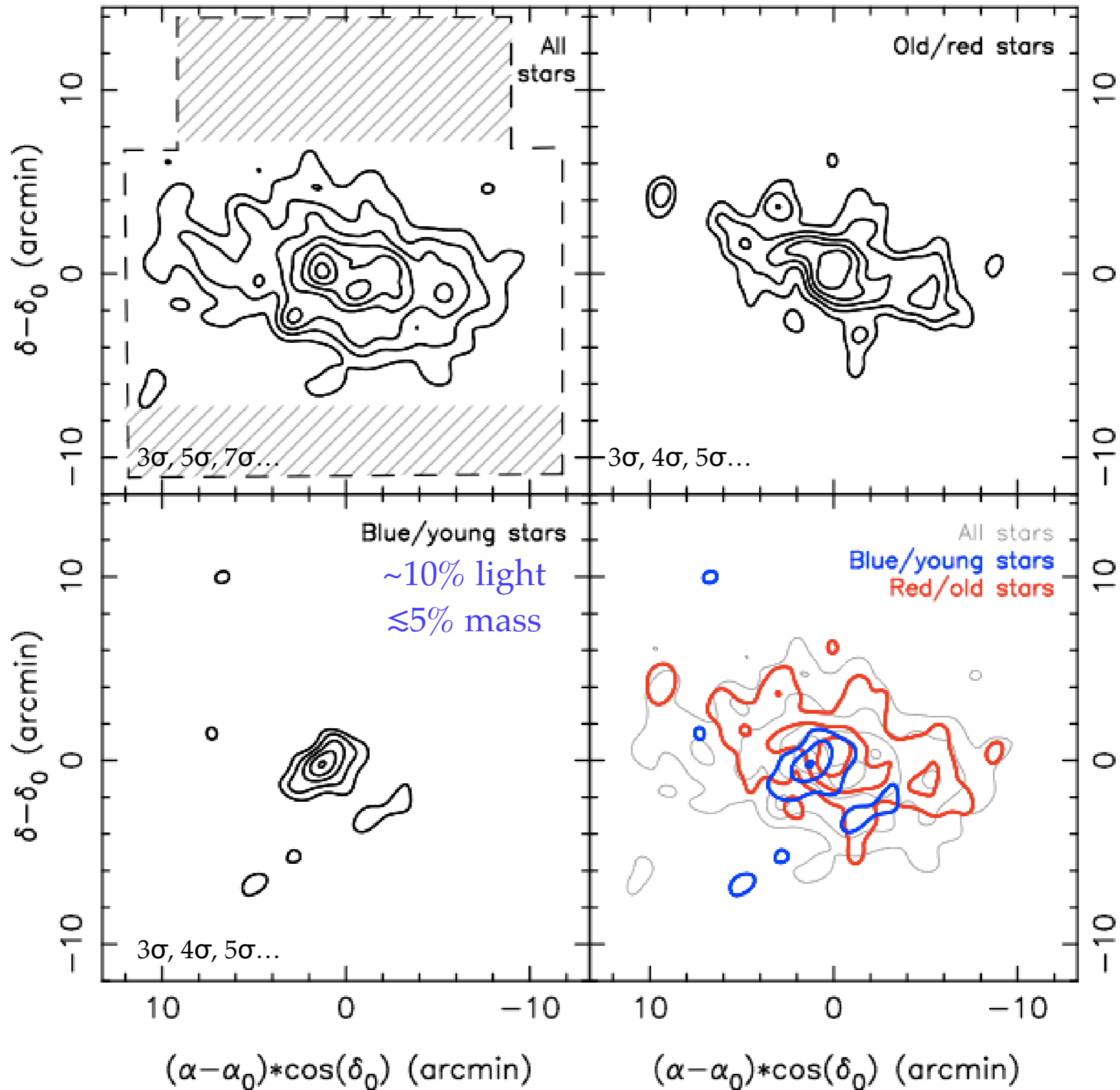


CVnI – the LBT view

Martin et al. (2007a)

within r_{hb}



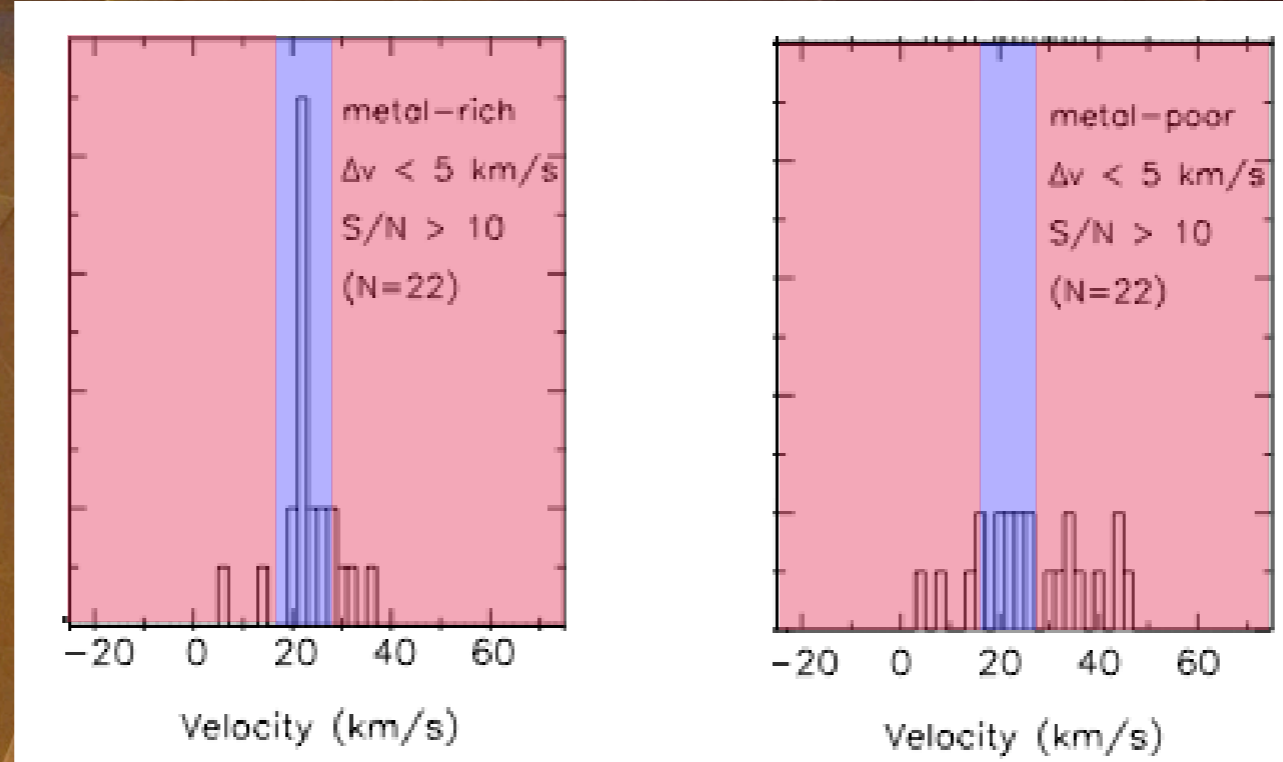


Canes Venatici I (Spectro)

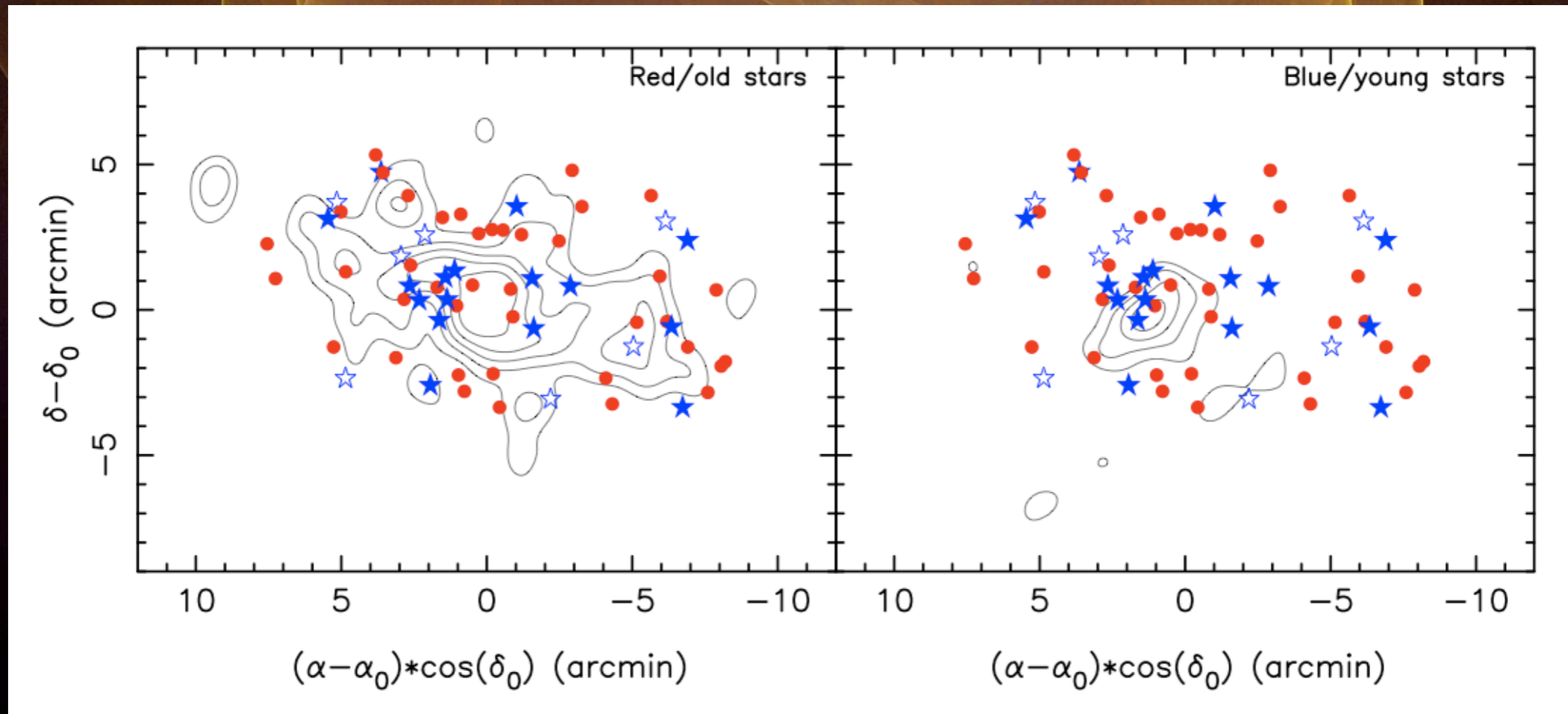
Ibata et al. (2006)

Martin et al. (2007a,b)

- Old >10 Gyr pop
- Metal-poor(er)
[Fe/H] ~ -2.0
- Extended
- Hot ($\sigma \sim 10$ km/s)

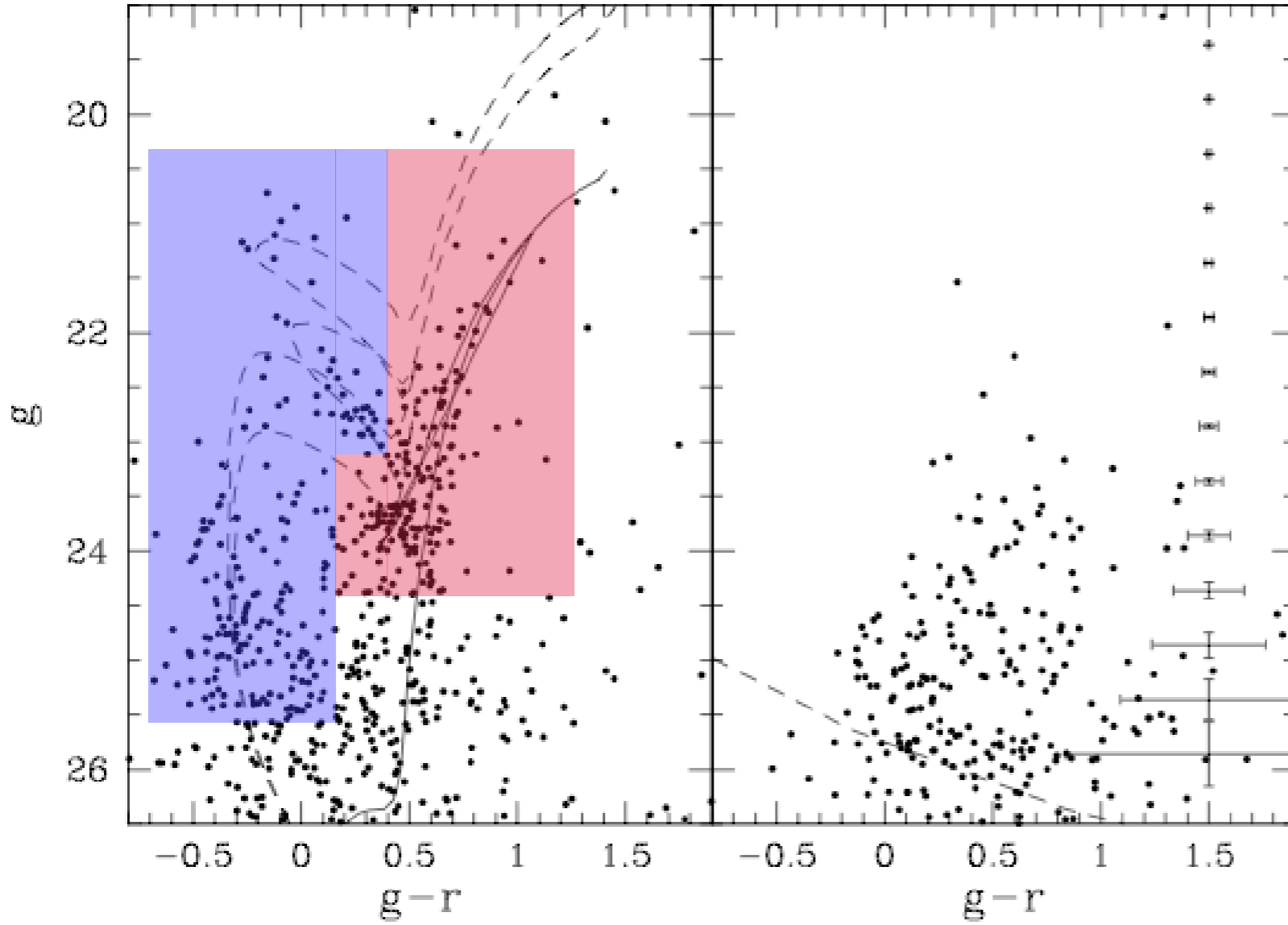


- Young ~ 2 Gyr pop
- Metal-rich(er)
[Fe/H] ~ -1.5
- Spatially confined
- Cold ($\sigma < 2$ km/s)



Leo T

de Jong et al. (2007b)



* Light:

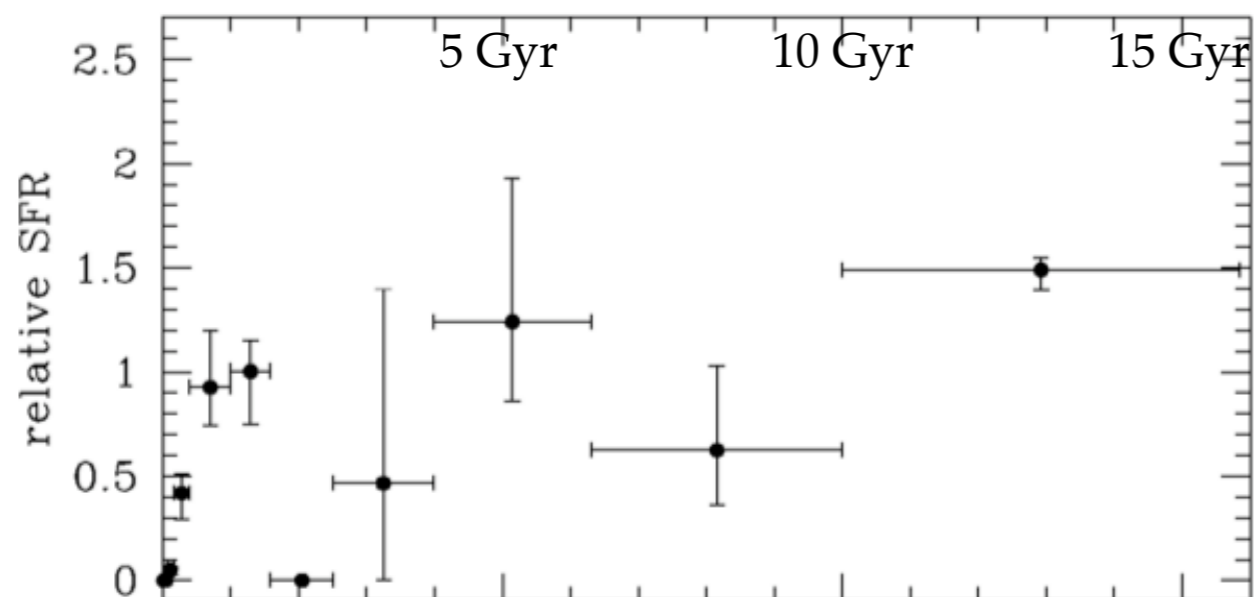
★ 65% old pop.

★ 35% young pop.

* Young pop less extended

★ young $r_{hb} = 104 \pm 8$ pc

★ old $r_{hb} = 148 \pm 16$ pc



Questions

* How can Hercules be so flat?

★ Disk-like?

- *but* $\sigma = 5.1$ km/s

★ Tidal distortion?

- seen in some dwarf galaxies (UMi, Sgr...) and tentative (2σ) detection of clumps in SDSS
- *but* $\sigma = 5.1$ km/s $\rightarrow R_{\text{peri}} \sim 8$ kpc \rightarrow very elliptical orbit (especially with $v_{\text{r,gsr}} = 145$ km/s)

★ Triaxial halo?

* CVnI & Leo T look pretty much like brighter galaxies (young to intermediate stars)

★ Is there a minimum mass to retain gas?

- Leo T: $0.8 \times 10^7 M_{\odot}$, CVnI: $2.7 \times 10^7 M_{\odot}$ (Simon & Geha mass estimates)

★ Cold pop in CVnI?

- star cluster formed 2Gyr ago? survival time?
- ~~is there still gas in CVnI?~~ $\rightarrow < 3 \times 10^4 M_{\odot}$